

# Forecasting Photosynthetic Photon Flux Density Under Cloud Effects: Novel Predictive Model Using Convolutional Neural Network Integrated With Long Short-term Memory Network

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## Research Article

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1       **Forecasting photosynthetic photon flux density under cloud effects: *novel***  
2       ***predictive model using convolutional neural network integrated with long short-***  
3       ***term memory network***

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14      **Abstract**

15      Forecast models of solar radiation incorporating cloud effects are useful tools to evaluate the impact  
16      of stochastic behaviour of cloud movement, real-time integration of photovoltaic energy in power  
17      grids, skin cancer and eye disease risk minimisation through solar ultraviolet (UV) index prediction  
18      and bio-photosynthetic processes through the modelling of solar photosynthetic photon flux density  
19      (PPFD). This research has developed deep learning hybrid model (*i.e.*, CNN-LSTM) to factor in role  
20      of cloud effects integrating the merits of convolutional neural networks with long short-term memory  
21      networks to forecast near real-time (*i.e.*, 5-minute) *PPFD* in a sub-tropical region Queensland,  
22      Australia. The prescribed CLSTM model is trained with real-time sky images that depict stochastic  
23      cloud movements captured through a Total Sky Imager (TSI-440) utilising advanced sky image  
24      segmentation to reveal cloud chromatic features into their statistical values, and to purposely factor  
25      in the cloud variation to optimise the CLSTM model. The model, with its competing algorithms (*i.e.*,  
26      CNN, LSTM, deep neural network, extreme learning machine and multivariate adaptive regression  
27      spline), are trained with 17 distinct cloud cover inputs considering the chromaticity of red, blue, thin,

28 and opaque cloud statistics, supplemented by solar zenith angle (*SZA*) to predict short-term *PPFD*.  
29 The models developed with cloud inputs yield accurate results, outperforming the *SZA*-based models  
30 while the best testing performance is recorded by the objective method (*i.e.*, CLSTM) tested over a  
31 7-day measurement period. Specifically, CLSTM yields a testing performance with correlation  
32 coefficient  $r = 0.92$ , root mean square error  $RMSE = 210.31 \mu \text{ mol of photons m}^{-2} \text{ s}^{-1}$ , mean absolute  
33 error  $MAE = 150.24 \mu \text{ mol of photons m}^{-2} \text{ s}^{-1}$ , including a relative error of  $RRMSE = 24.92\%$   $MAPE$   
34  $= 38.01\%$ , and Nash Sutcliffe's coefficient  $E_{NS} = 0.85$ , and Legate & McCabe's Index  $LM = 0.68$   
35 using cloud cover in addition to the *SZA* as an input. The study shows the importance of cloud  
36 inclusion in forecasting solar radiation and evaluating the risk with practical implications in  
37 monitoring solar energy, greenhouses and high-value agricultural operations affected by stochastic  
38 behaviour of clouds. Additional methodological refinements such as retraining the CLSTM model  
39 for hourly and seasonal time scales may aid in the promotion of agricultural crop farming and  
40 environmental risk evaluation applications such as predicting the solar UV index and direct normal  
41 solar irradiance for renewable energy monitoring systems.

42 **Keywords** Photosynthetic radiation; deep learning; stochastic cloud effects; solar radiation  
43 modelling; photosynthetic photon flux density; risk evaluation

## 44 **1.0 Introduction**

45 The global solar radiation used by plants in photosynthesis spans about 400-700 nm wavelength,  
46 which is a relatively narrow part of the entire solar spectrum, but one containing only about half the  
47 solar energy. Within this limits can be defined both the energy available for photosynthesis, the  
48 Photosynthetically Active Radiation (*PAR*,  $\text{Wm}^{-2}$ ) or alternatively, the Photosynthetic Photon Flux  
49 Density (*PPFD*;  $\mu \text{ mol of photons m}^{-2}\text{s}^{-1}$ ) [1] that will now be the subject of this paper. Lipid proteins,  
50 forming the building block of terrestrial and marine food webs, contribute to global biomass derived  
51 from agricultural animal and plant products that continue to be a growing source of worldwide energy  
52 production. Currently, green biofuels account for 11% of the world's total energy supply [2] coming

53 from primary plant and vegetable oil crops, secondary lignocellulosic by-products [3, 4], and third  
54 generation, enriched lipid microalgae bioproducts.

55 Significant research has focused on the optimisation of biofuel production particularly  
56 through the efficient production of microalgae photo-bioreactors (PBR) that can optimise the light,  
57 temperature, nutrient loads, and continuity of microalgae species [5-7]. Recent research works  
58 concentrated on the genetic modification of microalgae species for optimal acclimation to the  
59 environment. These are aimed at enhancing the overall output efficiency of the targeted microalgae  
60 products [8-10]. Alternative energy resources for PBR have also been investigated by including  
61 artificial light or organic fluorescent dyes to maximise solar conversion into optimal photosynthetic  
62 radiation bands [4]. Costs of artificial light sources have to date restricted the development of PBRs  
63 that do not retain enough access to reliable sources of photosynthetic-active solar radiation.  
64 Importantly, the availability of open-air setups utilising natural sunlight continues to be the most  
65 economically viable solution to farm microalgae and develop sustainable bio-products. These systems  
66 are by far the most prevalent, roughly occupying 90% of all third-generation commercial biofuel  
67 production facilities [11]. They are however dependent on both long and short-term fluctuations in  
68 localized-scale solar radiation where production can be improved by monitoring farms with robust  
69 forecasting efforts especially in real-time scales.

70 Solar radiation, affected by season, latitude and temporal variations in cloud cover, ozone,  
71 and atmospheric aerosols, influences the optimal utilisation of light at any given biomass production  
72 system, including its effect on plant growth or overall health. Typically, tropical environments that  
73 produce consistently high levels of solar insolation at the earth's surface are ideal [12]. However  
74 tropical climates are frequently affected by strong seasonal precipitation patterns resulting in  
75 fluctuations in solar light intensity. Cloud cover alone can drop the available Photosynthetic Photon  
76 Flux density (*PPFD*), which can reach 2000  $\mu\text{mol}$  of photons  $\text{m}^{-2} \text{s}^{-1}$  at noon, by as much as 80% [5,  
77 13]. Broken cloud can bring about short-term cloud enhancement of solar radiation (up to ~20%) and  
78 such conditions can bring about rapid fluctuation of solar radiation both above and below the clear

79 sky values. Yet, ideally, efficient biomass production requires a steady and reliable supply and  
80 monitoring of *PPFD* [15].

81 As net primary productivity is strongly influenced by climatic factors, much effort has been  
82 expended on measuring (and subsequently monitoring) the *PPFD*. A review of literature shows some  
83 limitations in terms of current predictive approaches where most methods have used monitoring rather  
84 than real-time forecasting approaches. Remote sensing platforms have been used to determine  
85 vegetation net production efficiency [16] and as a result can be used to determine the best locations  
86 for establishing farms, greenhouses or other high value agricultural hubs [12, 17-19]. Satellite remote  
87 sensing methods inherently must approximate the geometric absorption, scattering and transmission  
88 of clouds from relatively low resolution single-direction reflectance [20]. The most important  
89 environmental predictors to determine the global *PPFD* on the earth's surface are the annual  
90 precipitation, monthly cloud fraction, bioclimate layer information and month [21-25].

91 Having identified the best location for crops, the next step would be to forecast solar radiation  
92 conditions so that crops are protected and their growth is optimised. The seasonal and climatic factors  
93 which can be readily sourced from public datasets have been employed in previous AI-based  
94 approaches too, particularly to accurately predict agricultural crop yield, drought indices and rainfall  
95 in Pakistan [28, 29], China [30], USA [31] and Australia [32-34]. Such AI-based approaches are  
96 becoming useful tools to derive agricultural and biomass product efficiency mapping on a much  
97 broader scale where accurate surface instrumentation and local climate records are not available.  
98 Hemispherical photographs have been used to estimate *PPFD* with limited success[35]. Another  
99 approach has been artificial neural network (ANN) models that map out the available global surface  
100 *PPFD* using remote satellite products as predictor variables. This model, however, is based on an  
101 ANN approach that requires environmental predictors to produce an accurate forecast system [36].

102 Biomass productivity is not only dependent on total *PPFD* but also the diffuse fraction of  
103 *PPFD* [37]. Methods for retrospective *PPFD* estimation employ a mixture of remote satellite  
104 products, global reanalysis of climate information [38] and local surface instrumentation [39] to

105 model both direct and diffuse photosynthetic-active radiation and output biomass for a range of  
106 ecological and agricultural applications have also been developed [21].

107 In respect to solar energy, monitoring or integration into electricity grids, intermittencies in  
108 power production are highly driven by cloud variations [40]. However, the ability to develop reliable  
109 models to predict short-term (*e.g.*, 5-10 minute) solar radiation can provide a future solar system real-  
110 time monitoring capability to resolve clean energy challenges by better capturing cloud cover,  
111 lifetime, spread or stochastic movements. Also, the option to capture cloud cover variations in a solar  
112 ultraviolet index (UV Index) model such as the one developed previously by Deo *et al.*, [41] can help  
113 in skin cancer and eye disease risk mitigation. Developing a PPFD prediction model trained with  
114 cloud images may provide useful insights into UV index, solar power production or energy demand  
115 monitoring.

116 In a previous study, the near real-time PPFD prediction model of Deo *et al.* [39] was based  
117 on an adaptive neuro-fuzzy inference system to predict *PPFD* over 5-minute horizons in Queensland  
118 (Australia), using time lagged *SZA* data under cloud-free conditions. Utilising the local solar zenith  
119 angle (*SZA*) as the only input variable, they demonstrated good accuracy in predicting the real-time  
120 *PPFD* with changes in *SZA* for 5 minute and hourly forecasts. Such studies that model real-time solar  
121 photosynthetic energy can play a pivotal role in helping explore regional development of the  
122 agricultural sector. However, the inclusion of cloud cover (which is vital for the control of plant  
123 growth, was not considered in previous studies. The development of an AI-based model to predict  
124 the influence of cloud variations at near real-time, and how the cloud properties (derived from image  
125 chromic information) might control the amount of ground-based photosynthetic-active radiation is  
126 yet to be explored.

127 This paper develops an artificial intelligence (AI)-approach that considers the total sky  
128 conditions, addressing the role of cloud cover variations to accurately model *PPFD* at 5-minute time  
129 scales. The contribution and novelty are to build a first deep learning AI method for real-time *PPFD*  
130 forecasting, capturing the influence of cloud properties on measured photosynthetic-active radiation.

131 A deep learning-based methodology utilising whole sky image characteristics of both the cloud and  
132 cloud-free conditions typical to local farming environments incorporates data features from high  
133 temporal resolution images such as those captured by Total Sky Imager (TSI) or geo-stationary  
134 satellites *e.g.*, Himawari 8 or 9 providing inter-minute level sky images. The objectives are as follows.  
135 (1) To process *TSI*-based cloud images corresponding to *PPFD* measured at 5-minute intervals  
136 through a custom-built cloud segmentation algorithm [42] applied to each image, and produce  
137 descriptive statistics based on the blue, red, thin and opaque cloud chromatic features (*i.e.*, means,  
138 standard deviations, differences, ratios). These are then used to build an optimal set of model inputs  
139 (*i.e.*, cloud image properties) against a target (*i.e.*, *PPFD*). (2) To develop deep learning-based  
140 convolutional neural network and long short-term memory network (CLSTM) model following our  
141 earlier study [43], implemented for near real-time *PPFD* forecasting. (3) To benchmark the CLSTM  
142 model *w.r.t* conventional machine learning (MARS, ELM) and deep learning LSTM, CNN and DNN  
143 methods tested on the same training and testing subsets. To pursue the objectives, the present study  
144 has utilised data from a local TSI as a proof of concept. The parameters employed are cloud fraction,  
145 cloud type and the red-green-blue cloud chromatic properties derived from segmented sky images,  
146 with respect to simultaneous *PPFD* measurement at the subtropical location of Toowoomba (27.6°S),  
147 Australia.

## 148 **2.0 Theoretical Overview**

149 The theoretical details of deep learning (*i.e.*, CNN, LSTM, DNN) and conventional machine learning  
150 (ELM and MARS) methods are described elsewhere [43-48]). The CLSTM model, constructed by  
151 integrating CNN and LSTM, had been used elsewhere in natural language processing where emotions  
152 were analysed with text inputs [49], in speech processing where voice search tasks were performed  
153 using CLDNN combining CNN, LSTM and DNN [50], in video processing with CNN and Bi-  
154 directional LSTM models built to recognize human actions in video sequences [51], in the medical  
155 area where the CNN-LSTM method was developed to detect arrhythmias in electrocardiograms [52]  
156 and in industrial areas where a convolutional bi-directional LSTM model was designed to predict tool



157 wearing [53]. Other studies with CLSTM are evident, for example, time series application for  
158 prediction of residential energy consumption [54] [55], solar radiation prediction [43, 56-58] and  
159 wind speed prediction [59-61] as well as stock market applications in the prediction of share prices  
160 [62, 63]. In the solar radiation forecasting area, the study of Ghimire *et al.* [43] has developed a  
161 CLSTM model and compared its performance against the CNN, LSTM and DNN-based models,  
162 showing that the CLSTM model outperformed the standalone version of both CNN and LSTM  
163 models.

164 <Fig. 1>

165 Following earlier implementations [43], in this study we integrate CNN and LSTM to produce  
166 a hybrid system that ensures most prevalent data features are extracted using CNN prior to the  
167 sequential modelling of real-time photosynthetic radiation at 5-minute intervals. This objective model  
168 is depicted by a simplified schematic architecture in Figure 1. Generally, a CNN system is known to  
169 extract local trends or other features as well as common features recurring in time series at different  
170 intervals [64] and then used to serve as further inputs to LSTM model's architecture. LSTM is able  
171 to capture both the short- and the long-term dependencies in data patterns (*e.g.*, linking PPFD  
172 variability against time-based cloud movements) to learn the time sequential relationships among  
173 predictors and a target [65, 66]. First introduced for object recognition in image processing [67], the  
174 CNN model has a prominent structure composed of many convolution layers, pooling layers and one  
175 or more fully connected layer [62]. The primary building block applies a convolution filter (*i.e.*, a  
176 kernel function) for input data to generate a feature mapping scheme [68]. Using different filters,  
177 many sets of convolutions are performed in order to create different feature maps [69]. These are  
178 eventually combined to produce the convolution layer's final output. In the pooling layer, each feature  
179 map's dimension is reduced through down-sampling thereby mitigating the risks of model overfitting  
180 and reducing the model's training time [70]. The fully-connected layer at the end of the CNN is  
181 replaced with LSTM via the flattening layer to produce the hybrid CLSTM predictive model [71].

182 Other than the CLSTM model, the present study has utilised a standalone LSTM as a variation  
183 on Recurrent Neural Network (RNN) composed of memory cells coupled through layers, rather than  
184 the neurons in a conventional ANN-type model [72]. The RNN is generally considered to be  
185 somewhat incompetent in describing long-term dependences due to the gradient vanishing  
186 phenomenon [73]. Because of this, LSTM was developed by Hochreiter and Schmidhuber in 1997  
187 [74] and enhanced by Graves in 2013 [75]. In contrast to the classic RNN where gradients back-  
188 propagate exponentially, the LSTM model allows for gradients to flow unchanged by employing a  
189 cell memory. By using input gate, a forget gate, and an output gate, the LSTM unit can decide what  
190 to remember and what to forget and is therefore capable of addressing long-term dependencies. [76].  
191 In general, an LSTM block is made of the sigmoid ( $\sigma$ ) and hyperbolic tangent ( $\tanh$ ) layers, and two  
192 operations including pointwise summation ( $\oplus$ ) and multiplication ( $\otimes$ ) operations, as shown  
193 schematically in Figure 1. Mathematically, these processes can be defined by equations 1-6 [43].

194 Input gate  $i_t$ :

$$195 \quad i_t = \sigma(w_i x_t + R_i h_{t-1} + b_i) \quad \langle 1 \rangle$$

196 Forget gate  $f_t$ :

$$197 \quad f_t = \sigma(w_f x_t + R_f h_{t-1} + b_f) \quad \langle 2 \rangle$$

198 Output gate  $y_t$ :

$$199 \quad y_t = \sigma(w_y x_t + R_y h_{t-1} + b_y) \quad \langle 3 \rangle$$

200 Cell  $c_t$ :

$$201 \quad c_t = f_t c_{t-1} + i_t \bar{c}_t \quad \langle 4 \rangle$$

$$202 \quad \bar{c}_t = \sigma(w_c x_t + R_c h_{t-1} + b_c) \quad \langle 5 \rangle$$

$$203 \quad \text{Output vector } h_t: h_t = y_t \sigma(c_t) \quad \langle 6 \rangle$$

204 where,  $\sigma$  and  $\tanh$  are activation functions in the range [0,1] and [1,1] respectively,

$$205 \quad \text{Sigmoid function: } \sigma(\gamma) = \frac{1}{1+e^{-\gamma}} \quad \langle 7 \rangle$$

206 Hyperbolic-tangent function:  $\sigma(\gamma) = \frac{e^\gamma - e^{-\gamma}}{e^\gamma + e^{-\gamma}}$ . <8>

207  $b_i, b_f, b_y$  denote the input, forget, and output gate bias vectors, respectively;

208  $c_{t-1}$  and  $h_{t-1}$  are the previous cell and its output vector;

209  $h_t$  is the output vector;

210  $x_t$  denotes the input vector;

211  $w_i, w_f,$  and  $w_y$  are the matrix of weights from the input, forget, and output gates to the input,  
212 respectively; and

213  $R_i, R_f,$  and  $R_y$  define the matrix of weights from the input, forget, and output gates to the input,  
214 respectively.

## 215 **3.0 Materials and Method**

### 216 **3.1 Experimental Apparatus and Data Acquisition System**

217 Photosynthetic photon flux density, *PPFD*, was measured with corresponding cloud cover images at  
218 the Toowoomba Campus of The University of Southern Queensland 120 km west of Brisbane,  
219 Australia. Fig. 2(a) shows the geographic location of the study site. At the University's Atmospheric  
220 and Solar Ultraviolet Radiation Laboratory, a quality-controlled monitoring station measured *PPFD*  
221 and weather conditions since 2011 (Fig. 2b). Located at an elevation of 690 m *ASL*, Toowoomba is a  
222 regional city with a high solar energy potential and is also classified as a regional centre for  
223 agricultural activities that makes the *PPFD* forecast models an advantageous tool for practical  
224 applications in agricultural sectors. The specific study site also has a relatively large number of full  
225 sunshine days and a clear hemispheric view of the solar horizon [77] that also makes it an ideal site  
226 to implement the CLSTM model for real-time forecasting of photosynthetic-active radiation.

227 <Fig 2(a-d)>

228 To build the proposed CLSTM predictive model, high-quality, yet cloud-influenced  
229 measurements of *PPFD* were acquired over the austral summer solstice period (01 to 31 Mar 2013).  
230 The data were collected using a Quantum sensor (LI-190R; LI-COR, Lincoln, USA) connected to a  
231 CR100 Campbell Scientific data logger (Logan, USA) (Fig. 2). The LI-190R automated system was  
232 installed on an unobstructed rooftop site to continuously monitor the photosynthetic-active radiation  
233 at 5-minute intervals over a 24-hr period. Employed in several other research works [39, 78, 79], the  
234 LI-190R system is mainly designed for long-term, outdoor usage with a manufacturer-stated  
235 uncertainty of  $\pm 5\%$  traceable to the US National Institute of Standards and Technology. In this paper,  
236 the *PPFD* time series for the daytime period 07.00 AM—05.00 PM were used, considering that solar  
237 irradiance is mainly intercepted by plants during daytime, and that the night level of photosynthetic  
238 energy is practically zero.

239 **<Fig 3(a-b)>**

240 Figure 3(a) shows the temporal patterns in measured *PPFD* time series sampled at 5-minute  
241 intervals, ranging from 0 to 2300  $\mu\text{ mol of photons m}^{-2}\text{ s}^{-1}$  but this variation over entire diurnal cycles  
242 is different for different days or times. This is perhaps due to cloud cover or atmospheric conditions  
243 (*e.g.*, ozone, aerosols, water vapor). Fig. 3(b) shows a sample of five cloud images with their  
244 respective *PPFD* and solar zenith angle. It is noticeable that even for a similar value of *SZA* (28-29°)  
245 at 10.55 AM (10 Mar) and 12.55 PM (15 Mar), the value of *PPFD* varies by almost 28%. Similar  
246 observation can be made for the data on 01 March (06.55 AM) and 30 March (16.55 PM) measuring  
247 the *PPFD* values of 54  $\mu\text{ mol of photons m}^{-2}\text{ s}^{-1}$  and 333  $\mu\text{ mol of photons m}^{-2}\text{ s}^{-1}$ . Meanwhile here is  
248 rather similar *PPFD* for March 30<sup>th</sup> and March 5<sup>th</sup> even though *SZA* changes considerably. This  
249 illustrates how cloud fraction is an important modulator of *SZA*-controlled photosynthetic-active  
250 radiation, including cloud height and depth that are not considered in this analysis.

251 **3.2 Sky Image Processing and Cloud Segmentation**

252 A quick and efficient self-adaptive Python-based tool called the *TSI Analyser* developed in earlier  
253 work [42] is employed for sky image segmentation and extraction of cloud chromatic properties from  
254 images obtained by Total Sky Imager (*TSI*) instrument (serial number: 175). Details of the *TSI*  
255 *Analyser* algorithm are described elsewhere [42] but in principle, it is able to produce cloud cover-  
256 based statistical properties for *every* image that is associated with a measured *PPFD* value. This aims  
257 to capture the overall sky conditions, particularly, to include the contributory role of cloud cover  
258 variations in training the proposed CLSTM predictive model. To do this, we refer to comparisons  
259 between red and blue intensities in clouds, red-blue ratios, and red-blue difference. We also  
260 segmented each image into the normalized red-blue-ratio that was undertaken in our earlier paper [36]  
261 based on the commonly used red-blue ratio [80] such that the *TSI440*-based pixel values of each of  
262 the red and blue channels were determined. It is noteworthy that the normalized ratios are consistent  
263 with conventional cloud detection methods with practical importance in cloud segmentation [81]. It  
264 is also important to note that the red (*R*) to the blue (*B*) ratio maintains a higher relative resolution  
265 despite the down sampling that occurs when the images are saved in *.jpeg* format. To acquire images,  
266 the *TSI440* enables a user defined threshold for opaque and thin clouds [82] with the latter cloud type  
267 presenting a difficulty in cloud segmentation especially when aerosols are present [83], which is not  
268 further considered in this study, assuming everything captured by the user threshold to be thin cloud.

269 The *TSI Analyser* was applied to a 1-month dataset with 5-minute interval cloud images  
270 considering over 200,000 images collected at a  $480 \times 320$  spatial resolution. These whole-sky images  
271 have been captured using *TSI440* [84-86] used in previous research (*e.g.*, [82, 87, 88]). The *TSI440*  
272 instrument consists of a reflective dome with a camera suspended above it [89, 90] pointing  
273 downwards to generate a *.jpeg* format colour image of the whole sky. A non-corrupted sky image  
274 array is then read using commands from the *NumPy* library in Python [91] by means of the OpenCV  
275 library's *imread* command [92]. This is converted from OpenCV's blue-green-red (*BGR*) to red-  
276 green-blue (*RGB*) format for further image processing.

277 Table 1 summarises the data for cloud chromatic properties derived from segmented images  
278 including the descriptive statistics (*i.e.*, mean, standard deviation, difference, & ratio) based on the  
279 blue, red, thin, and opaque cloud (pixelized) features *per* image. The segmentation algorithm  
280 produced the average of the whole sky blue ( $B_{av}$ ), whole sky red ( $R_{av}$ ), as well as the statistical features  
281 based on standard deviation, ratios, or differences of the blue ( $B$ ) and red ( $R$ ) pixel values for clouds  
282 that represent the estimated proportion of pixelized cloud features likely to be a function of the  
283 photosynthetic-active radiation received at a measuring sensor. To analyse the degree of associations  
284 between cloud movement and an instantly measured *PPFD* value, a cross correlation analysis is  
285 performed to determine the covariance measured by  $r_{cross}$  prior to developing the proposed CLSTM  
286 model. Table 1 includes the  $r_{cross}$  used to determine the order of our model input combinations,  
287 presented in Table 2. It is evident that the average of whole sky-blue pixel in a total sky image appears  
288 to generate the largest value of  $r_{cross} \sim -0.747$ , followed by the standard deviation of the blue cloud  
289 pixel ( $r_{cross} \sim -0.640$ ). This exceeds an  $r_{cross}$  value of  $-0.631$  computed for solar zenith angle that is  
290 traditionally used as the only predictor variable of photosynthetic-active radiation as per other studies  
291 (*e.g.*, [39]). This analysis also shows that the covariance of the whole sky-blue average and the  
292 standard deviation of the blue cloud pixels are more strongly correlated with *PPFD* compared with  
293 the *SZA* dataset.

294

### <Table 1>

295 To corroborate the findings in Table 1 we now inspect visually the covariance in cloud chromatic  
296 properties against measured photosynthetic-active radiation. Figure 4 displays a scatterplot of the  
297 cloud cover statistics as well as *SZA* data that are regressed against the measured *PPFD* in the model  
298 training phase. The whole sky-blue average is seen to attain the highest coefficient of determination  
299 ( $r^2 = 0.549$ ) with respect to the *PPFD* values. The other significant predictor variables are found to  
300 be the blue cloud pixel standard deviation ( $r^2 = 0.403$ ), solar zenith angle ( $r^2 = 0.403$ ) and the standard  
301 deviation of the whole sky-blue ( $r^2 = 0.365$ ). It is especially notable that the ratio of red to blue sky  
302 and the difference between the blue and red pixels in a whole sky image appears to be weakly

303 correlated with *PPFD* data series, and therefore, may not contribute significantly towards improving  
304 the proposed CLSTM model. Taken together, the present analyses clearly ascertain that at least two  
305 of the cloud chromatic properties (*i.e.*, whole sky blue & blue cloud pixel averages associated with  
306 measured *PPFD*) are more strongly correlated with *PPFD*, compared with the solar zenith angle used  
307 in earlier studies. This deduction confirms that the inclusion of cloud cover properties may be a crucial  
308 task used to improve earlier models for photosynthetic-active radiation (*e.g.*, [39]).

309

#### <Fig 4>

310 A comparison of the *PPFD* data series within the first 7 days of model training data is made  
311 against cloud-image derived predictor series in Figure 5. Note that here, the first 847 points are  
312 employed to demonstrate the association of *PPFD* and cloud property before developing the proposed  
313 CLSTM predictive model. While the changes in *PPFD* are not well-represented by *SZA* due to the  
314 solar zenith angle presenting a much smoother variation over any given diurnal cycle, there is a clear  
315 temporal correspondance between the magnitude of *PPFD* with many of the cloud-image statistical  
316 features. This correspondance is especially pronounced on the *x*-axis scale from the datum point 363  
317 to 847 for image pixels representing the whole sky blue average and its standard deviation, and the  
318 standard deviation of the blue cloud pixels. Interestingly, for the whole sky red average pixels, the  
319 standard deviation of the red cloud pixels, the average of blue cloud pixels, the whole sky red-blue  
320 ratio, the standard deviation of the whole sky red and the difference of red-blue pixels are also  
321 demonstrating a good degree of harmony in terms of their temporal variation against the *PPFD*  
322 timeseries. While the direct association between some of the cloud chromatic properties is not so  
323 clear, as expected, there does appear to be a moderating effect in terms of the jumps in *PPFD* against  
324 any cloud property. This indicates that the subtle, yet non-linear effects of cloud movements on  
325 photosynthetic-active radiation should be captured in a *PPDF* forecast model.

326

#### <Fig 5>

327

#### <Table 2>

### 328 3.3 Predictive Model Design

329 To develop the objective hybrid model (*i.e.*, CLSTM) and benchmark (or comparative) models using  
330 deep learning (LSTM, CNN, DNN) and machine learning (ELM & MARS) algorithms, both the  
331 python [93] and the MATLAB-based [94] scripts were implemented on Intel *i7* computer with 3.40  
332 GHz processor running on 32GB memory. Figure 6 illustrates the model development stage and Table  
333 2 lists the input combinations used in all designated models together with the details of data  
334 partitioned in the training (53.3%), validation (23.3%), and testing (23.3%) subsets.

335 <Fig. 6>

336 To build an accurate CLSTM model that can consider the role of cloud cover variations,  
337 particularly by using cloud chromatic properties to generate near real-time photosynthetic-active  
338 radiation forecasts, an optimal arrangement of the model's inputs is firstly deduced. A sequential  
339 ordering approach (*e.g.*, [95]) is adopted where ranked cross-correlation coefficients  $r_{\text{cross}}$  deduced  
340 from the respective predictor variable as illustrated Table 1 (*i.e.*, cloud-based time series, or solar  
341 zenith angle derived from an empirical method [96]). This proposed method led to the first predictive  
342 model ( $M_1$ ) being constructed using the average of whole sky blue ( $B_{\text{av}}$ ) pixels, followed by the second  
343 model ( $M_2$ ) with both the  $B_{\text{av}}$  and the standard deviation of blue cloud pixels ( $BC_{\text{sd}}$ ) pixels and the  
344 third model ( $M_3$ ) having  $B_{\text{av}}$ ,  $BC_{\text{sd}}$  and solar zenith angle ( $SZA$ ) as enunciated by Table 2.

345 By inclusion of cloud properties, this study advances earlier work [39, 88] where  $SZA$  was the  
346 only predictor used to forecast  $PPFD$  and solar UV index ignoring cloud variations. This study  
347 advances the standard approaches [39, 88] that utilize only  $SZA$  neglecting the role of clouds in  
348 modulating  $PPFD$ . It is noteworthy that successive addition of series based on  $r_{\text{cross}}$  concurs with  
349 earlier prediction problems [95] aimed at evaluating potential improvements in CLSTM model. To  
350 evaluate the utility of a cloud-free model, a standard approach used in photosynthetic-active radiation  
351 [39], solar UV index [88] and global solar models [95]), a CLSTM model designated as  $M_{18}$ , with



352 only the *SZA*, was constructed as a reference model without any inclusion of cloud cover properties.  
353 Overall, the model design process resulted in 18 distinct predictive models, as stated Table 2.

354 <Fig. 7>

355 As this study's intent is to build a forecast model that can accurately predict the  
356 photosynthetic-active radiation at a future timescale over near real-time (5-minute) intervals, we have  
357 further explored the cross correlation between cloud chromatic properties and photosynthetic-active  
358 radiation (or *PPFD*) using a time-lagged correlogram. Figure 7 identifies the covariance between  
359 *PPDF* (*i.e.*, target) and *SZA*, along with all of the other cloud-image derived predictor variable data  
360 in the model training phase. Evidently, the lagged series show a strong ( $\pm$ ) serial correlation exceeding  
361 the statistically significant region at the 95% confidence which is indicated by a blue line.  
362 Interestingly, the correlation coefficient in terms of the time-shifted cloud properties for non-zero lag  
363 (*i.e.*, occurring for an input that was regressed on a target at a different timescale) is also prominent  
364 for some of the inputs (*e.g.*, thick clouds, average of red pixel values in the cloud cover, difference  
365 between whole sky red and the blue pixels, and the ratio of red to the blue pixel values in the clouds).  
366 This indicates a strong non-linear association between cloud chromatic properties and photosynthetic-  
367 active radiation, potentially indicating the need for a non-linear modelling approach to forecast  
368 photosynthetic-active radiation. To construct the proposed CLSTM model, all of the cloud chromatic  
369 properties and the *SZA* measured over a time lag of 5 minutes is used:

$$370 \quad PAR(t + 1) = f\{X(t)\} \quad (9)$$

371 where  $PAR(t + 1)$  denotes the photosynthetic photon flux density (*PPFD*,  $\mu\text{mol of photons m}^{-1}\text{s}^{-1}$ ) at  
372 a next time interval of 5-minute time horizon,  $X(t)$  is the relevant input and  $t$  is the time scale. Prior  
373 to the modelling process, all inputs and the target were scaled to be between [0, 1] where:

$$374 \quad X_N = \frac{X - \hat{X}}{\hat{X} - \bar{X}} \quad (10)$$

375 where

376  $X_N$  = Normalized values of a variable  $X$

377  $X$  = Actual value of a variable  $X$

378  $\hat{X}$  = Maximum value of a variable  $X$

379  $\check{X}$  = Minimum value of a variable  $X$

380 To identify the contributory effects of cloud variations in forecasting 5-minute  
381 photosynthetic-active radiation, this study firstly develops a 3-layered convolutional neural network  
382 (CNN) and long short term memory network (LSTM) with a 4-layered deep neural network (DNN),  
383 and multivariate regression spline (MARS) and extreme learning machine (ELM) models. Following  
384 the benchmark methods, CNN and LSTM algorithms were integrated in accordance with earlier study  
385 [43] to generate a 4-layered objective model (denoted as hybrid CLSTM). For model development  
386 parameters, see Appendix (Table A1 a-c). In general, for the CLSTM architecture, the first half  
387 comprised of the CNN used for feature extraction whereas the second half comprised of the LSTM  
388 algorithm used to forecast *PPFD* by incorporating these CNN-grained input features.

### 389 **3.3.1 Common Hyperparameters for Deep Learning (DL) Models**

390 Open-source DL Python libraries, Scikit-Learn [97] and Keras[98, 99] were used to implement CNN,  
391 LSTM and DNN algorithms. Hyperparameters of all benchmark models were deduced through grid  
392 search. In this study, the DL models share the following four common hyperparameters.

- 393 • *Activation functions*: Except for the output layer, all of the network layers relied on the same  
394 activation function, which accords to the other studies [100, 101] so we have used the rectified  
395 linear unit (*ReLU*) [102].
- 396 • *Dropout*: This is considered as a potential regularization to minimize overfitting issues in  
397 order to improve the training performance [103]. The dropout aims to select a fraction of the  
398 neurons (defined as a real hyperparameter over the range 0 to 1) at each model iteration and  
399 prevent them from retraining [104-106]. For this study, this fraction of neurons was  
400 maintained to be 0.1.

- 401 • *Two statistic regularization.* This included L1 (*i.e.*, least absolute deviation) and L2 (*i.e.*, least  
402 square error) applied together with the dropout. It is imperative to mention that the role of L1  
403 and L2 penalization type parameters is to minimize the sum of the absolute differences and  
404 the sum of the square of the differences between the forecasted and target PPF values,  
405 respectively [107-109]. Also, the addition of a regularization to the loss is to encourage  
406 smooth network mapping in the DL network, particularly by penalizing the large parameters  
407 values to reduce the level of nonlinearity in the network models [110, 111].
- 408 • *Early stopping:* The issue of overfitting can be further addressed by introducing an early  
409 stopping (ES) phase in Kera [98, 112] so that the mode is set to a minimum while the patience  
410 is set to 30 [110, 113, 114]. This is done to also ensure that the training process will terminate  
411 when the decrease in the validation loss has stopped for a number of patience-specified epochs  
412 [115-117].

### 413 **3.2.2 CNN Hyperparameters and Hybrid CNN-LSTM Model Development**

414 The CNN model's hyperparameters were also optimised that included the following options.

- 415 • *Filter size:* The size of the convolution operation filter was optimised.
- 416 • *Number of convolutions:* The number of convolutional layers in each CNN was optimised.
- 417 • *Padding:* This study has utilized the same padding in order to ensure that the input feature  
418 map and output feature map dimensions were identical [118].
- 419 • *Pool-size:* A pooling layer was used between each convolution layer to avoid further  
420 overfitting. This pooling layer also helps decrease the number of parameters and network  
421 complexity [119]. In this study, we have utilized a pool-size of 2 between the layer 1 and 2 of  
422 the CNN model.

423 Finally, the hybrid CNN-LSTM model comprised of 3 convolutional layers, with pooling operations  
424 where a selection of the convolutional layer channels was based on grid search process. In the model's

425 architecture, the outputs of flattening layer served as the inputs of LSTM recurrent layer while the  
426 LSTM recurrent layer was directly linked to the final output.

### 427 **3.4 Non-deep Learning Benchmark Models**

428 This study develops ELM and MARS models (as benchmark methods) considering their relative  
429 success in solar predictive problems [88]. The ELM architecture composes of a single hidden layer  
430 system with 17 input neurons (to enable cloud cover and *SZA*-based inputs to be fed in) (Table 3c), a  
431 maximum of 1000 hidden neurons and 1 output neuron allocated to the forecasted *PPFD*. To optimise  
432 the ELM model, this study tests several activation functions (*i.e.*, sine, hard limit, radial basis,  
433 triangular basis, logarithmic sigmoid & tangent sigmoid equations) following earlier approach [88]  
434 with an optimal model achieved using logarithmic sigmoid equation indicated in Table 3(c). To  
435 identify an optimal ELM architecture, the hidden neuron was varied from 1 to 1000 with each  
436 architecture then evaluated on a validation dataset (25% in this study) to identify the optimal  
437 architecture. As ELM requires random initialization of hidden layer parameters, the model was run  
438 1000 times with the lowest root mean square error (*RMSE*) over all hidden nodes used to select the  
439 optimal ELM model. The optimal ELM was denoted as 10–23–1 (input–hidden–output) which  
440 included 10 predictor variables and 23 hidden neurons to attain the most accurate forecasts of *PPFD*  
441 data.

442 For the MARS model, an ARESLab-based MATLAB toolbox (ver. 1.13.0) [120] is adopted.  
443 Out of the two basis functions (*i.e.*, cubic & linear) within its piecewise equation, the cubic form is  
444 adopted [121] given its capacity to handle multiple predictors. The generalized recursive partitioning  
445 regression (RPR) is also employed as an adaptive algorithm for function approximation [122] with  
446 the process including a forward and backward deletion process to reach the optimal MARS equation.  
447 In the forward phase, a ‘naïve’ model with just the intercept term is used with iterative addition of  
448 the reflected pair(s) of basis functions to generate the maximum decrease in the model training error  
449 based on *RMSE*. the model with the lowest Generalized Cross-Validation statistic was selected.

450 Table 3(c) also lists the optimal MARS model equation. For greater details about ELM and MARS,  
 451 readers can consult earlier References [88].

### 452 3.5 Predictive Model Performance Evaluation

453 The study adopts the model performance metrics recommended by American Society for Civil  
 454 Engineers [123] to evaluate the hybrid CLSTM (and all the other benchmark) models. By appraising  
 455 the degree of agreement between  $PPFD_{for}$  and  $PPFD_{obs}$  the computed metrics include correlation  
 456 coefficient ( $r$ ), mean absolute error ( $MAE$ ,  $\text{mmol m}^{-2}\text{s}^{-1}$ ), root mean square error ( $RMSE$ ,  $\text{mmol m}^{-2}\text{s}^{-1}$ ),  
 457 including the relative % magnitudes of  $RMSE$  and  $MAE$ , Legate & McCabe's ( $LM$ ) and the Nash  
 458 Sutcliffe's coefficient ( $E_{NS}$ ). Mathematically, these are as follows [43, 124-126]:

$$459 \quad r = \left( \frac{\sum_{i=1}^N (PPFD_{for,i} - \overline{PPFD}_{obs,i})(PPFD_{for,i} - \overline{PPFD}_{obs,i})}{\sqrt{\sum_{i=1}^N (PPFD_{for,i} - \overline{PPFD}_{obs,i})^2} \sqrt{\sum_{i=1}^N (PPFD_{for,i} - \overline{PPFD}_{obs,i})^2}} \right) \quad <11>$$

$$460 \quad MAE = \frac{1}{N} \sum_{i=1}^N |(PPFD_{for,i} - PPFD_{obs,i})| \quad <12>$$

$$461 \quad RMSE = \sqrt{\frac{1}{N} \sum_{i=1}^N (PPFD_{for,i} - PPFD_{obs,i})^2} \quad <9>$$

$$462 \quad MAPE = \frac{1}{N} \sum_{i=1}^N \left| \frac{(PPFD_{for,i} - PPFD_{obs,i})}{PPFD_{obs,i}} \right| \times 100 \quad <14>$$

$$463 \quad RRMSE = \frac{\sqrt{\frac{1}{N} \sum_{i=1}^N (PPFD_{for,i} - PPFD_{obs,i})^2}}{\frac{1}{N} \sum_{i=1}^N (PPFD_{obs,i})} \times 100 \quad <15>$$

$$464 \quad LM = 1 - \left[ \frac{\sum_{i=1}^N |PPFD_{obs,i} - PPFD_{for,i}|}{\sum_{i=1}^N |PPFD_{obs,i} - \overline{PPFD}_{obs,i}|} \right], \quad 0 \leq LM \leq 1 \quad <16>$$

$$465 \quad E_{NS} = 1 - \left[ \frac{\sum_{i=1}^N (PPFD_{obs,i} - PPFD_{for,i})^2}{\sum_{i=1}^N (PPFD_{obs,i} - \overline{PPFD}_{obs,i})^2} \right], \quad -\infty \leq E_{NS} \leq 1 \quad <10>$$

466 where  $PPFD_{obs}$  and  $PPFD_{for}$  are the observed and forecasted  $i^{th}$  value in test period,  $\overline{PPFD}_{obs}$  and  
467  $\overline{PPFD}_{for}$  are the observed and forecasted means and  $N$  is the number of datum points within a test set.

468 The present study adopts several performance measures for a robust evaluation of the forecast  
469 models specially to overcome the constraints of any single metric. Diagnostic tools and graphical  
470 representations utilising scatterplots and error distribution are used in conjunction with statistical  
471 indices to test the versatility of 5-minute forecasts models.

#### 472 **4.0 Results and Discussion**

473 In this section the results generated by the hybrid CLSTM predictive model, including the other deep  
474 learning-based (LSTM, CNN, DNN) and machine learning-based (ELM, MARS) models are  
475 appraised by checking the degree of congruence between measured and forecasted photosynthetic-  
476 active radiation at a 5-minute temporal scale. A careful evaluation of the results emanating from the  
477 cloud cover-based models using various input combinations (*i.e.*, Table 2) and a reference model  
478 utilising only the solar zenith angle is also made, to identify the contributory role of cloud variations  
479 in modelling photosynthetic photon flux density ( $PPFD$ ). Figure 8 shows a scatterplot of the tested  
480 data where the performance of CLSTM (and comparative models) is evaluated in terms of the degree  
481 of agreement between observed and forecasted  $PPFD$ . Also included are the results of deep learning-  
482 based LSTM, CNN and DNN, as well as the other machine learning-based (MARS & ELM) model.  
483 Note that in here, only the optimally trained model (out of the 17 designated input combinations,  
484 Table 2) considering the influence of cloud variations on 5-minute  $PPFD$ , are shown.

485 While the performance of the newly proposed CLSTM model seems to exceed that of the  
486 other predictive models, as evidenced by the largest  $r^2$  (~0.846), the gradient (representing the  
487 forecasted and observed  $PPFD$ ) closest to unity, and the smallest bias constant, it also had a capped  
488 maximum forecasted  $PPFD$ . (Fig. 8a), the most accurate prediction differs significantly for the  
489 different model types and their input combinations. For example, the best performance of the CLSTM  
490 model (Fig. 8a) is attained through  $M_8$ :  $PAR = f\{B_{av}, BC_{sd}, SZA, B_{sd}, OC, R_{av}, RC_{sd}, BC_{av}\}$ . This

491 means that the CLSTM model requires cloud segmented properties based on the whole sky blue  
 492 average, standard deviation of the blue pixels, blue cloud average pixels, standard deviation of the  
 493 blue cloud pixels, opaque cloud pixels, standard deviation of the red cloud pixels, whole sky red  
 494 average pixels, and the *SZA* time series yielded the most accurate performance. For the case of the  
 495 LSTM model (Fig. 8b), the best performance is attained through  $M_{13}$ :  
 496  $PAR = f \left\{ B_{av}, BC_{sd}, SZA, B_{sd}, OC, R_{av}, RC_{sd}, BC_{av}, \frac{R_{av}}{B_{av}}, R_{sd}, RBC_{diff}, BRC_{diff}, TC \right\}$  with this  
 497 model using the eight input variables that are already used in CLSTM as well as the time series of  
 498  $\frac{R_{av}}{B_{av}}, R_{sd}, RBC_{diff}, BRC_{diff}$  and  $TC$  to generate the best performance. A similar deduction is made for  
 499 CNN, ELM and MARS models where the designated model  $M_{11}$ ,  $M_{10}$  and  $M_{12}$  is seen to generate the  
 500 highest coefficient of determination compared with a lower  $r^2$  value for the other input combinations  
 501 specified in Table 2. When the best input combination for the DNN model is deduced by progressively  
 502 adding the cloud cover properties one by one, the model  $M_5$  generates the best performance ( $r^2 =$   
 503  $0.810$ ) with an input combination  $PAR = f \{ B_{av}, BC_{sd}, SZA, B_{sd}, OC \}$ . Note that in this case, only five  
 504 input series (*i.e.*, whole sky-blue average and standard deviation of whole sky blue including the  
 505 standard deviation of blue cloud pixels, solar zenith angle, and opaque clouds) are required. However,  
 506 it is also noteworthy that the performance of the DNN model is relatively lower than CLSTM model  
 507 (*i.e.*,  $r^2 = 0.810$  vs.  $0.846$ ). The analysis reveals that, while the hybrid CLSTM model integrating the  
 508 LSTM and CNN methods used to emulate 5-minute *PPFD* far exceeds the performance of all other  
 509 comparative models, their inputs combinations (based on cloud properties and *SZA*) appear to be  
 510 unique indicating the different capabilities for feature extraction required to accurately predict the  
 511 photosynthetic-active radiation.

512 <Figure 8>

513 In congruence with previous results shown in Figure 8, the frequency of the absolute value of  
 514 predicted error distribution in the testing phase generated by the *optimal* CLSTM and the *optimal*  
 515 benchmark models, are shown in Figure 9. It is notable the newly proposed CLSTM model (*i.e.*,  $M_8$ )

516 generated almost 75% of all predictive errors within the smallest error bracket *i.e.*,  $\pm 200 \mu \text{ mol}$  of  
517 photons  $\text{m}^{-2} \text{ s}^{-1}$  band compared with LSTM,  $M_{13}$  (~72%), DNN,  $M_5$  (~69%), CNN,  $M_{11}$  (~69%), ELM,  
518  $M_{10}$  (~71%) and MARS,  $M_{12}$  (~63%). The largest frequency of predictive errors within the smallest  
519 error bracket no doubt concurs with a smaller frequency of redistributed forecast errors, albeit within  
520 a larger error band exceeding  $\pm 200 \mu \text{ mol m}^{-1} \text{ s}^{-1}$ . For example, we note that ~17% of all predictive  
521 errors attained by CLSTM are located within the  $\pm(200\text{--}400) \mu \text{ mol of photons m}^{-2} \text{ s}^{-1}$  whereas those  
522 for LSTM, DNN, CNN, ELM and MARS are seen to record ~21%, 22%, 21%, 20% and 27% of all  
523 predictive errors, respectively.

524

### <Figure 9>

525 Next, we investigate the overall statistical score metrics computed over the last 7 days of  
526 tested data (*i.e.*, 24-03-2013 to 31-03-2013) using 5-minute *PPFD*. Table 3 presents both the optimal  
527 model developed using various input combinations ( $M_1\text{--}M_{17}$ ), as well as the reference model ( $M_{18}$ )  
528 developed using traditional approach (*i.e.*, solar zenith angle only) as per earlier studies [39].  
529 Interestingly, the best performance among all tested models is attained by different input  
530 combinations that use both the cloud cover properties and the solar zenith angle as an input variable.  
531 However, for the predictive models developed with only the solar zenith angle as an input, the  
532 performance of all the deep learning (CLSTM, CNN, DNN, LSTM) and machine learning (ELM,  
533 MARS) models appear to be significantly inferior to those that utilise cloud cover properties and *SZA*.  
534 In fact, the *SZA*-based models produce the smallest magnitude of  $r$  (between 0.796–0.623), and the  
535 largest *RMSE* / *MAE* between 412.77–438.99 / 354.29 – 368.09  $\mu \text{ mol of photons m}^{-2} \text{ s}^{-1}$  within the  
536 testing phase. This contrasts the values  $r$  (0.894–0.920) and between 210.31–241.26  $\mu \text{ mol of photons}$   
537  $\text{m}^{-2} \text{ s}^{-1}$  for *RMSE* and 150.24–183.11  $\mu \text{ mol of photons m}^{-2} \text{ s}^{-1}$  for *MAE* for the models that incorporate  
538 cloud cover variations. This result indicates the important contributory role played by cloud cover  
539 variations in modulating the photosynthetic-active radiation and particularly, in improving the  
540 forecasting performance of the hybrid CLSTM and all of the other comparative models.

541

### <Table 3>



542 In Table 3, we also present several metrics for models developed using cloud cover as well as  
543 the *SZA* data where the normalised performance metrics based on the relative percentage error, Nash  
544 Sutcliffe coefficient, and the Legates & McCabe's Index is incorporated. It is noteworthy that the  
545 inclusion of cloud cover properties is seen to lead to an improved performance of the hybrid CLSTM,  
546 and all the other predictive models. That is, we note the smaller error values ranging between 24.92–  
547 28.79% (*RRMSE*) and 38.01–56.21% (*RMAE*) for models utilising cloud cover properties, whereas  
548 the errors based on *SZA* as the only input variable are relatively larger, between 49.15–51.98%  
549 (*RRMSE*) and 128.39–176.72% (*RMAE*). It is therefore deducible that appropriate factoring of the  
550 role of cloud cover variations to predict 5-minute *PPFD* can help reduce the forecasted errors very  
551 significantly. This deduction also concurs with a much higher value of the Nash-Sutcliffe and the  
552 Legate's & McCabe's Index obtained for all models that are trained with cloud cover properties. If  
553 the performance of only the hybrid CLSTM model is evaluated against all the comparative models;  
554 after factoring the cloud cover properties, we register the values of  $E_{NS}$  and  $LM$  to be 0.846 and 0.679  
555 compared with 0.796–0.829, and 0.607–0.660 for the case of ELM, LSTM, CNN, DNN and MARS  
556 models. Again, these metrics ascertain the influence of cloud cover on ground level photosynthetic-  
557 active radiation, and the superiority of the newly proposed CLSTM model.

558 <Figure 10>

559 Figure 10 is a Taylor diagram that evaluates all predictive models, including those with cloud  
560 cover properties and *SZA*-only as inputs. In this figure the most *optimal* model based on the best input  
561 combinations are compared to provide a visual framework for the forecasted *PPFD* against a  
562 reference (observed *PPFD*) data point. The pertinent statistics in Taylor diagram show the weighted  
563 centred pattern correlations and the ratio of the normalized root-mean-square (RMS) difference  
564 between the 'tested' data (*i.e.*, CLSTM, CNN, LSTM, DNN, ELM & MARS) and the 'reference'  
565 (observed) data. Two important deductions are made: *firstly*, it is clear that all of the *SZA*-based  
566 reference models are clustered much further away from the axis representing the observed *PPFD*  
567 whose *RMS*-centred difference certainly separates them away from the cloud cover-based models,

568 and *secondly*, the CLSTM model utilising cloud properties (indicated in red) is at a closest location  
569 to the observed *PPFD*, and also attains the highest correlation among all tested predictive models. It  
570 is also observable that all the cloud cover-based models are within a smaller cluster (and hence,  
571 demonstrate comparable performance) whereas those utilising *SZA* only are more scattered. This  
572 suggests that the inclusion of cloud cover is necessary to optimise all the DL and ML models, but  
573 among all these models, the CLSTM remains the superior choice to forecast the 5-minute *PPDF*  
574 dataset.

575

### <Figure 11>

576 In Figure 11, we investigate the nature of the predictive error generated by the objective model  
577 (*i.e.*, CLSTM) and the counterpart models while also evaluating the role of cloud cover variations  
578 using the modelled *PPDF* from the *SZA* only, and the cloud cover-based predictive models. Here, the  
579 forecast error  $|FE| = |PPFD_i^{for} - PPFD_i^{obs}|$  is illustrated as a boxplot for both the cloud property-  
580 based and the *SZA*-based model. There is a clear consensus that the best model out of the ones  
581 designated as  $M_1$ – $M_{17}$  utilising cloud features as inputs are able to attain a significantly lower error  
582 distribution compared to the reference model  $M_{18}$  where *SZA* is the only predictor variable. For all  
583 predictive models trained with the *SZA* input data, the maximum error value is manyfold higher, and  
584 so is the upper quartile, median and the lower quartile of  $|FE|$ . This means that when cloud feature is  
585 excluded from a predictive model the ability to forecast *PPFD* values is much less, and this can result  
586 in a wider distribution of the errors for the *SZA*-based model. A comparison of all models developed  
587 using cloud cover properties, including the *SZA*, certainly shows a much smaller lower quartile, upper  
588 quartile, maximum and median values of the forecasted error. When all models trained with cloud  
589 features are investigated, the boxplots show the smallest value of 5-number summary, with the  
590 minimum, maximum, lower quartile, upper quartile and medians occupying smaller magnitudes for  
591 the case of CLSTM compared with CNN, LSTM, DNN, MARS and ELM. This is congruent with  
592 earlier results (Figs. 8–10) to demonstrate the CLSTM model as being the optimal choice to emulate  
593 the near real-time photosynthetic active radiation over a 5-minute scale.

594

<Figure 12>

595 To further establish the veracity of the hybrid CLSTM model Figure 12 shows the empirical  
596 cumulative distribution function (*ECDF*) of the error encountered in forecasting the photosynthetic-  
597 active radiation in the testing phase. The *ECDF* clearly demarcates the important role of cloud cover  
598 variations against the standard approach utilising *SZA* as the only input variable. A clear separation  
599 point is noted throughout the *ECDF* such that all models trained with cloud cover inputs attain a much  
600 smaller forecasted error with a steeper rising curve in contrast to the slower growth in *ECDF* within  
601 larger error values. In fact, the cloud-property based models reach an asymptotic state around an  $|FE|$   
602 value of  $600 \mu \text{ mol of photons m}^{-2}\text{s}^{-1}$  whereas the *SZA*-based models continue to accumulate error  
603 values until  $|FE|$  values of  $900 \mu \text{ mol of photons m}^{-2}\text{s}^{-1}$ . Comparing the *ECDFs* of the hybrid CLSTM  
604 model against the other DL and ML models trained with cloud features, this result clearly concurs  
605 with Figure 9 where the growth in predictive errors is smaller for the CLSTM compared with the  
606 CNN, LSTM, DNN, ELM and MARS models. This establishes the efficacy of the newly developed  
607 CLSTM model trained with cloud cover features to generate the most accurate performance in terms  
608 of forecasting the 5-minute *PPFD* dataset.

609

<Figure 13>

610 We further explore the influence of cloud cover variations on the prescribed objective model  
611 (*i.e.*, CLSTM) in Figure 13 where the 5-minute forecasted *PPFD* valued averaged over the entire test  
612 dataset is shown *with* and *without* cloud cover input features. Note that these errors, showing both the  
613 percentage and absolute error values, are deduced from the forecasted and observed photosynthetic-  
614 active radiation measured from 07.00 AM to 05.00 PM. It is obvious that the hybrid CLSTM model  
615 utilising the cloud cover-based input features yields the smallest mean error over the whole diurnal  
616 cycle. The CLSTM error follow a temporal pattern where the models register relatively larger errors  
617 in early morning (~07.00 AM to 09.00 AM) and late afternoon (~04.00 PM to 05.00 PM) compared  
618 with the rest of the day. Possible causes for this error is that the CLSTM model did not isolate  
619 variability with solar zenith of clear sky aerosol optical thickness and cloud chromic properties

620 associated with forward and backscattering at the cloud edges [127-131] [x7] or aerosol. It is also  
621 possible that the CLSTM model is unable to capture enough features to predict the relatively smaller  
622 *PPFD* values in the morning and afternoon where the aerosol optical thickness is similar to the cloud  
623 scattering. Nonetheless, this analysis clearly outlines the important role of cloud cover conditions in  
624 modelling photosynthetic-active radiation and shows an important advancement in photosynthetic-  
625 active radiation prediction compared to earlier studies using the traditional (*SZA*) method.

## 626 **5.0 Further Discussion**

627 The results generated by the proposed CLSTM model have established relationships between  
628 photosynthetic-active radiation and cloud cover conditions necessary to model near real-time 5-  
629 minute *PPFD* with this objective model exhibiting the best performance against several other  
630 competing (*i.e.*, deep learning and machine learning-based) approaches. An incremental inclusion of  
631 cloud cover features based on time series of segmented cloud properties also captured a different, yet  
632 a significant contributory influence, further improving the testing performance of CLSTM model.  
633 However, improvements to the CLSTM model can be made with further development and refinement  
634 of the cloud segmentation tool itself.

635 The major contributions have led to significantly improved modelling approaches relative to  
636 earlier studies [132-135] where artificial intelligence models have utilised only the solar zenith angle,  
637 and failed to consider the effect of cloud cover conditions on photosynthetic-active radiation. Such  
638 methods used the more conventional modelling approaches (*i.e.*, single hidden layer neuronal  
639 architecture) without any deep mining of the predictive features as undertaken by the proposed  
640 CLSTM method in this paper. Given that the movement of clouds is highly variable depending on  
641 altitude and wind, cloud shape and thickness commonly vary on timescales of much less than 30  
642 minutes, our study has captured such influences on the ground-based photosynthetic active radiation  
643 at ~5-minutes. The modelling of photosynthetic radiation at this time interval is also of practical  
644 relevance in the monitoring and the supply of enough sunlight for solar energy generation or biofuels

645 exploration, monitoring the healthy growth of plants, monitoring day light integral or available  
646 photosynthetic energy for plant functions.

647 This pilot study has demonstrated how the CLSTM model utilising statistical input features  
648 from cloud images can become a sophisticated deep learning system for the future development of  
649 solar energy monitoring devices [136]. One such technology that can be particularly useful in the  
650 agricultural sector (*i.e.*, an automated monitoring and control system for algae photobioreactors) has  
651 practical relevance. For specific applications, CLSTM model can be incorporated into a smart  
652 environment monitoring system, 24 x 7, by adopting Internet of Things (IoT) and Wireless Sensor  
653 Networks, WSN [137] in a monitoring systems to ensure sustained health of crops and particularly  
654 considering how cloud conditions can affect their growth. The light available for microalgal  
655 photosynthesis remains a function of the surface solar irradiance over day-night cycles with  
656 environmental factors such as light, temperature, and nutrient status not only affecting photosynthesis  
657 and productivity of algae but also influencing the pattern, pathway and activities of cell metabolism  
658 or composition. Therefore, the efficacy of CLSTM model to forecast photosynthetic-active radiation  
659 at high temporal resolutions of 5-minutes that also matches a near real-time scale, can be trained on  
660 live cloud cover data or other atmospheric conditions. This application of the proposed deep learning  
661 system can help in regular prediction of the availability of sunlight in real time including its role in  
662 modelling temperature, water salinity, or nutrient status within an algae pond. The CLSTM model  
663 can also be employed in biophysical model platforms to improve the robustness of plant-growth  
664 models particularly, providing accurate estimations of photosynthetic photon flux density due to the  
665 scarcity of their ground-based measurements [138]. As the cost of Total Sky Imagers (TSIs) can be  
666 insurmountable for most solar energy or biofuel generation farm locations, geo-stationary satellites  
667 such Himawari 8 or 9, operating at roughly 10-minute interval and relatively high spatial resolutions  
668 may become good suppliers of sky images to be used as inputs for the CLSTM model to generate  
669 predicted PPFD or other components of solar radiation at appropriate temporal resolutions.

670 Other than agricultural applications, our CLSTM model incorporating cloud conditions also  
671 has potential use in public health and energy sectors. In an earlier study, Deo *et al.*, [88] developed a  
672 very short-term reactive system for solar ultraviolet (UV) prediction, albeit using a single hidden  
673 layer extreme learning machine (ELM) model and without any consideration to cloud cover  
674 conditions. Such a UV forecasting system can be a useful avenue for real-time prediction of UV  
675 radiation, a component of the solar spectrum known to cause melanoma and eye disease. However,  
676 as neither that study, nor any other prior or following study has incorporated the role of cloud cover  
677 conditions into a solar UV forecasting system, the proposed CLSTM system built on deep learning  
678 technology might be a viable tool to test the role of cloud conditions on UV prediction. One may  
679 therefore develop a CLSTM system for short-term (*e.g.*, 5-minute) reactive forecasting of UV index  
680 to help in public health risk mitigation. In terms of its application in energy industries, the CLSTM  
681 model can become a viable tool for real-time management of solar energy in a photovoltaic system  
682 by responding through a cloud image-based forecast system for solar power prediction, and  
683 particularly utilising cloud movements, cloud forms or its relative position-based features. Such a sky  
684 image-based solar power forecasting system utilising deep data mining can be of great value to the  
685 solar energy industry [40].

## 686 **6.0 Conclusions**

687 The industrial-scale production of solar power, biofuels and agriculture including food and health  
688 supplements from micro-algae farming, require reliably predicted solar radiation over short, long, and  
689 medium-term periods. This study has established the feasibility of predicting very short-term, 5-  
690 minute interval photosynthetic-active radiation using segmented cloud cover properties and solar  
691 zenith angle in a sub-tropical region in Toowoomba, Australia. A total of 17 different segmented  
692 cloud cover properties based on the mean, standard deviation, differences, and ratios of blue and red  
693 pixel values in clouds, including opaque and thin clouds (applied through thresholds on the Total Sky  
694 Imager), were acquired as part of the University of Southern Queensland Solar Radiation Monitoring  
695 Program running for more than 15 years. Together with the solar zenith angle, the cloud cover

696 properties based on segmented image inputs were applied to develop the hybrid deep learning (*i.e.*,  
697 CLSTM) model based on an integration of convolutional neural networks (to map out the cloud and  
698 SZA-based input features) and the long short-term memory network (to generate the near real-time  
699 forecasts of 5-minute photosynthetic photon flux density, *PPFD*). The CLSTM, verified to be highly  
700 superior in predicting 5-minute *PPFD* through 17 different predictor variable (or input) combinations,  
701 was benchmarked against three deep learning methods (*i.e.*, LSTM, CNN, DNN) and two machine  
702 learning (*i.e.*, ELM & MARS) methods. All these predictive models were evaluated using statistical  
703 score metrics and diagnostic plots visualising the degree of agreement between forecasted and  
704 observed photosynthetic photon flux density in an independent test dataset where the CLSTM model  
705 was applied.

706 The findings can be enumerated as follows.

707 (i) Among the objective (CLSTM) and five competing models, the best performance (out of 17  
708 distinct input combinations of segmented cloud properties) was attained by different  
709 combinations of cloud features. For example, the best CLSTM model  $M_8$  utilised average of  
710 whole sky-blue pixels, standard deviation of blue cloud pixels, *SZA*, standard deviation of the  
711 whole sky blue pixels, opaque clouds, averaged whole sky red pixels, standard deviation of  
712 red cloud pixels and the average of blue cloud pixels. By contrast, the second-best model (*i.e.*,  
713 ELM) used all the 8 inputs required by CLSTM, including two additional inputs (*i.e.*, ratio of  
714 whole sky blue to whole sky red average cloud pixels and whole sky red standard deviation)  
715 for its optimal model  $M_{10}$ . The third-best model, or LSTM required three additional inputs  
716 compared with ELM. The CNN model, which was the fourth-best model developed to forecast  
717 5-minute *PPFD* used only 11 input variables, whereas the DNN model relied on only 5 input  
718 variables. Despite different numbers of inputs used by the hybridised, deep learning and  
719 machine learning models, the performance of CLSTM remained superior.

720 (ii) In terms of comparing the *SZA*-only models, the CLSTM without cloud registered twice the  
721 model error (~50.07%) compared to with cloud ~24.92% in the testing phase. The other

722 metrics for SZA models only were also far less impressive for all models than those where  
723 clouds were incorporated. In terms of Taylor diagram comparing the different models to a  
724 reference (*i.e.*, observation) point, the non-cloud cover-based models were certainly scattered  
725 much further away from this reference point, and their performances were quite disparate  
726 relative to a comparable performance for cloud cover-based models (Fig. 10). Likewise, the  
727 distribution of forecast error was more widely spread, with significantly larger outliers, upper  
728 quartile, or extreme error values for SZA-only models (Figs. 11–12). These findings ascertain  
729 the important role of considering cloud cover variations to accurately model photosynthetic-  
730 active radiation.

731 Finally, this pilot study highlights the appropriateness of using cloud cover features to develop  
732 a deep learning method for very short-term, near real-time forecasting of photosynthetic-active  
733 radiation. If cloud segmented image properties from geo-stationary satellite images are available,  
734 the need for ground-based inputs that are data expensive for many regional locations can be  
735 eliminated. Furthermore, fish-eye lens or adapters used in mobile phones may also be able to supply  
736 the relevant images so the developed CLSTM model can be tried with those inputs to make the  
737 predictive model more accessible and applicable to all regions where the segmentation software is  
738 made available. This newly proposed method can offer major advantages in terms of the model  
739 implementation in regions with limited access to data such as agricultural farms. However, the present  
740 study only considers cloud properties using local, two-dimensional ground-based sky images so the  
741 inclusion of other atmospheric attenuations imposed by water vapour and aerosol should also be  
742 considered in the proposed CLSTM model with performance tested in different climatic zones and  
743 seasons. The improvement in CLSTM model's practical viability for other regions may also be made  
744 through its implementation on hourly, daily, and seasonal scales by sourcing satellite and other remote  
745 sensing products. Such testing of the proposed CLSTM predictive model in a wider range of climates,  
746 in both remote and regional locations is a necessary step to help in direct harnessing of solar energy,



747 biofuels from microalgae, agricultural crop monitoring and supporting bio-physical sectors where  
748 photosynthetic-active radiation needs to be monitored.

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# Figures

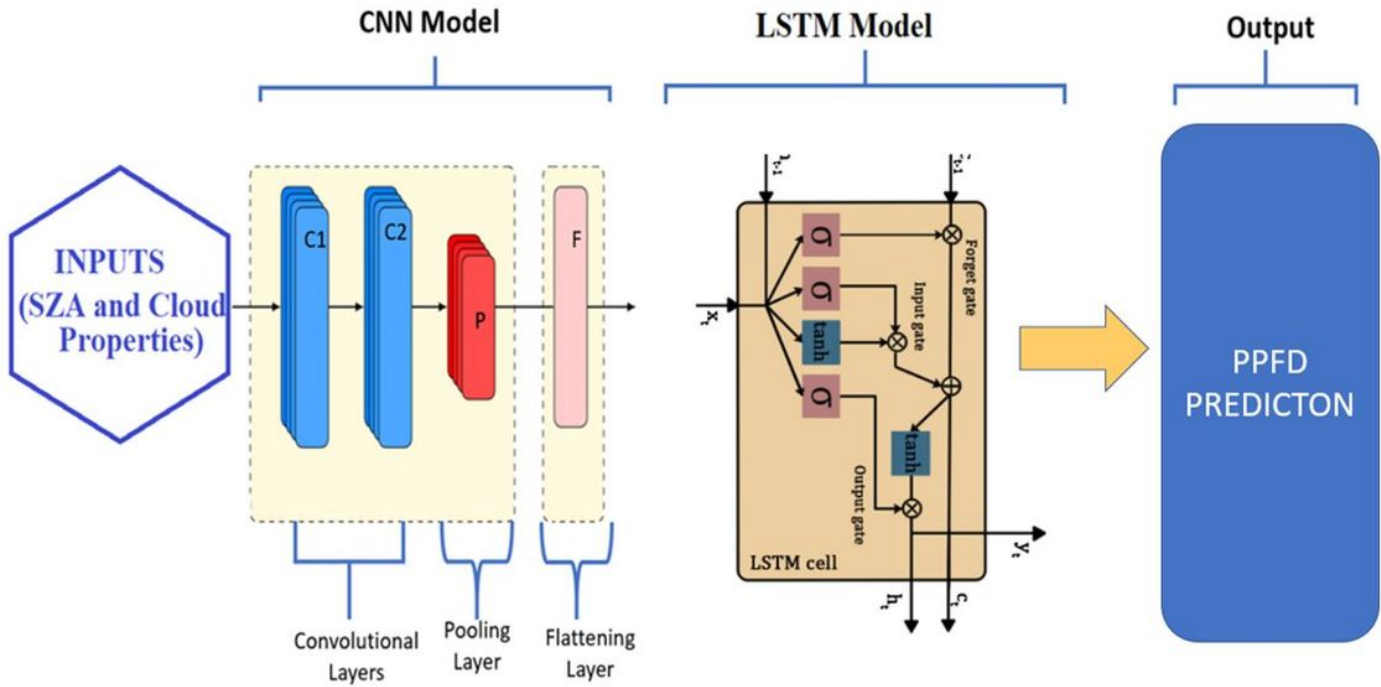
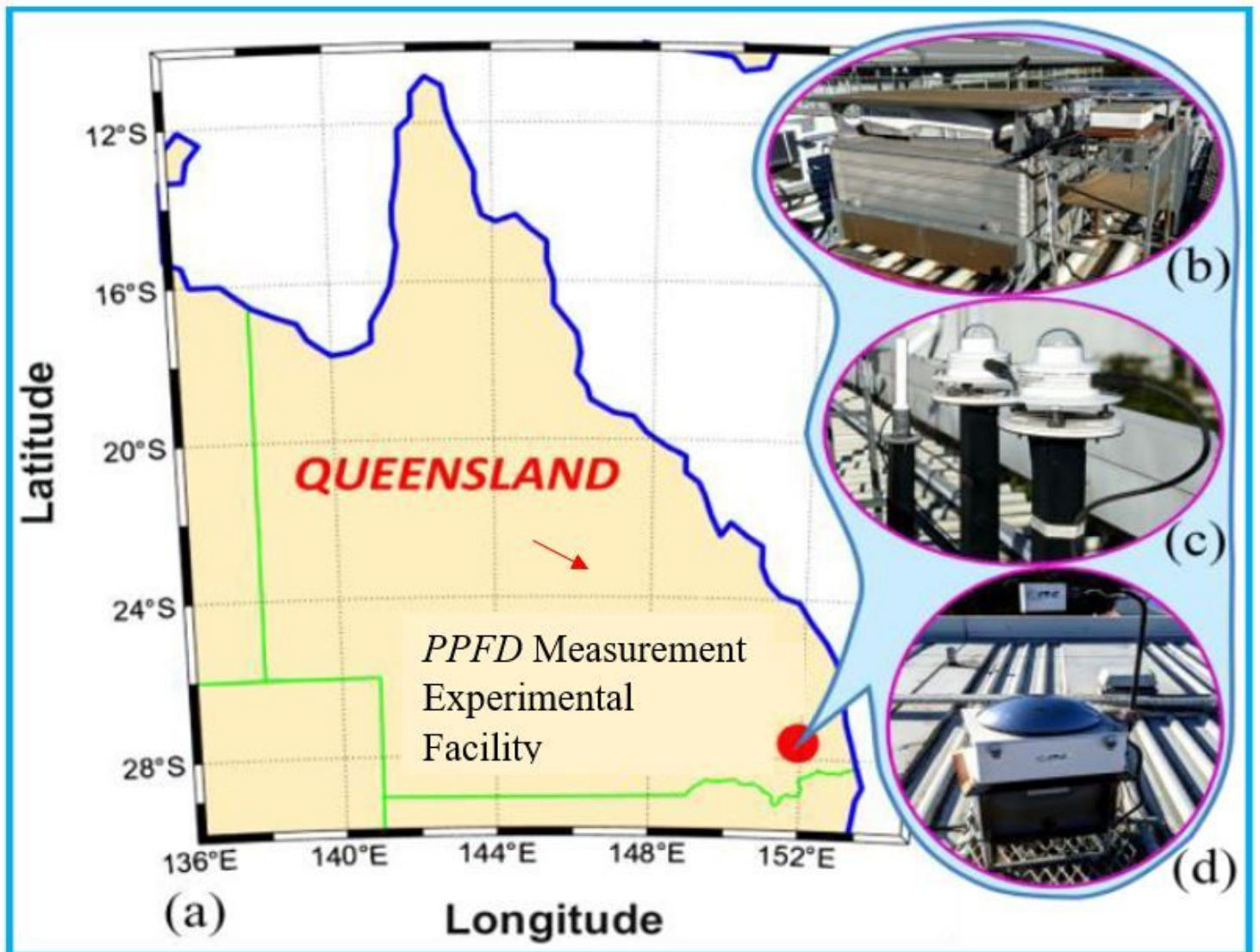


Figure 1

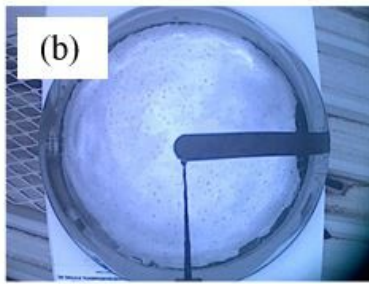
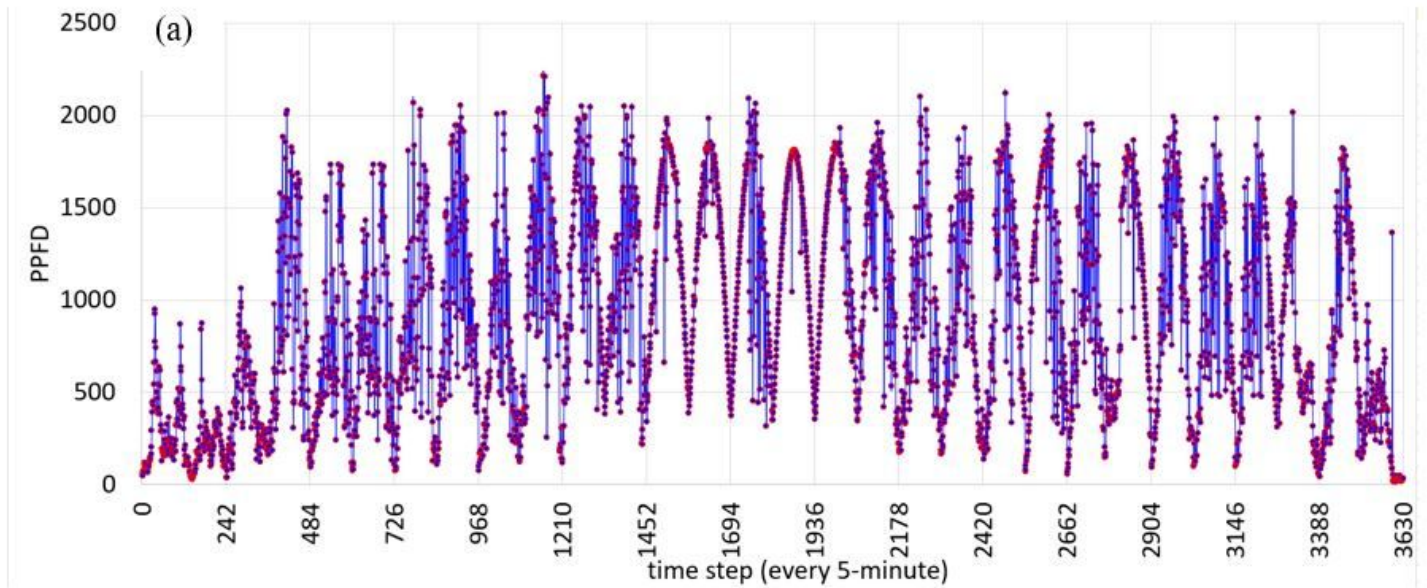
Schematic illustration of Convolutional Neural Network-Long Short-Term Memory Network (CLSTM) predictive framework. CNN used for feature extraction from solar zenith angle (SZA) and cloud chromatic properties from Total Sky Imager (TSI) and LSTM is used for time sequential modelling of the photosynthetic-active radiation (represented as photosynthetic photon flux density, PPFd).





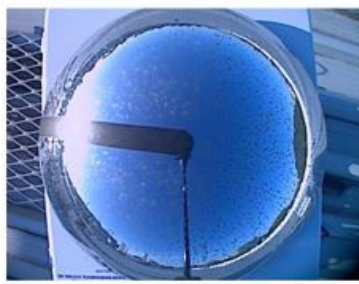
**Figure 2**

(a) Geographic location of the measurement facility in Queensland, Australia where CLSTM model is implemented. (b) Roof-top mounted Bentham DTM300 Spectroradiometer for 5-minute PPFD ( $\mu$  mol of photons  $m^{-2} s^{-1}$ ) measurement. (c) Co-located 501 broadband UVR Biometer. (d) Synchronous Total Sky Imager, TSI440 set-up to capture sky images and record solar zenith angle (SZA). Note that the LI-COR is connected to CR100 Campbell data logger at University of Southern Queensland Solar Research Laboratory.



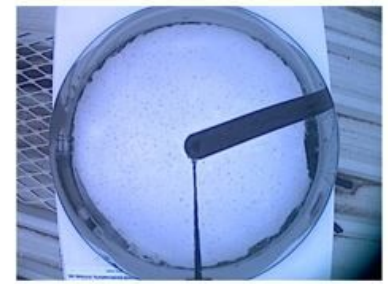
PPFD =  $54 \mu \text{ mol m}^{-2} \text{ s}^{-1}$  (low)

SZA =  $75^\circ$ : 1 Mar 06:55 AM



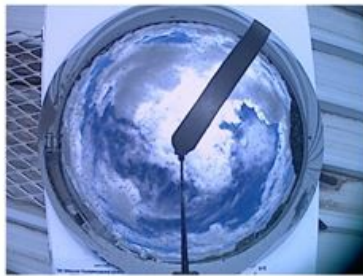
PPFD =  $333 \mu \text{ mol m}^{-2} \text{ s}^{-1}$  (medium)

SZA =  $78^\circ$ : 30 Mar, 16:55 PM



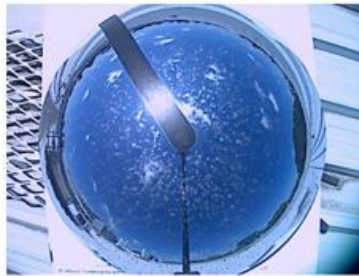
PPFD =  $410 \mu \text{ mol m}^{-2} \text{ s}^{-1}$  (medium)

SZA =  $50^\circ$ : 5 Mar, 08:55 AM



PPFD =  $1263 \mu \text{ mol m}^{-2} \text{ s}^{-1}$  (high)

SZA =  $28^\circ$ : 10 Mar 10:55 AM

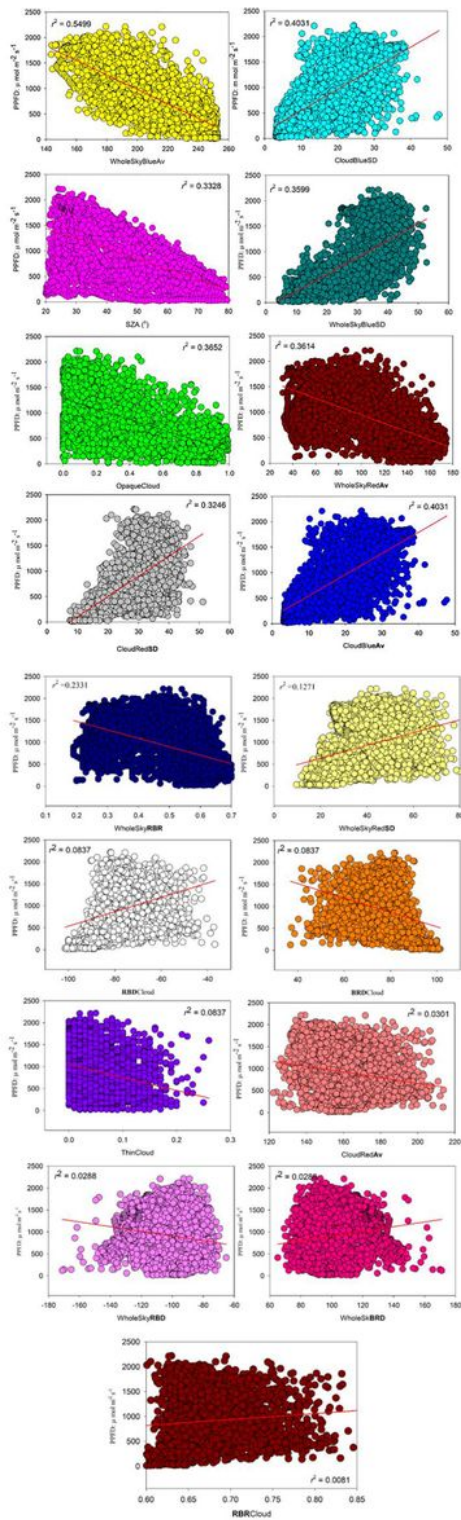


PPFD =  $1754 \mu \text{ mol m}^{-2} \text{ s}^{-1}$  (high)

SZA =  $29^\circ$ : 15 March, 12:55 PM

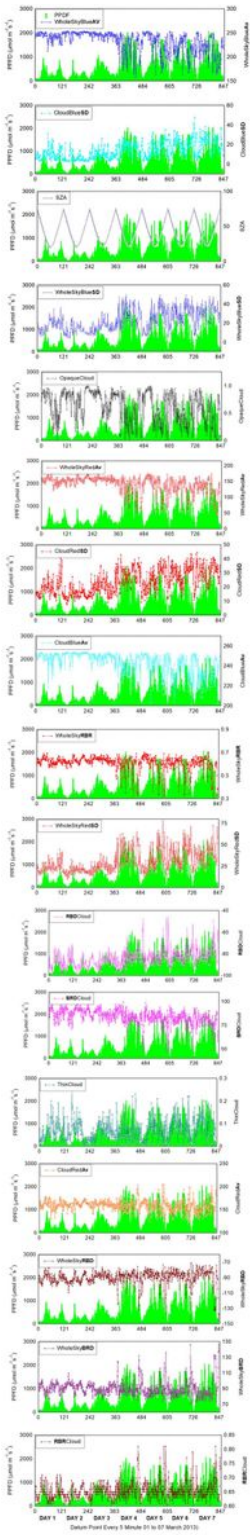
**Figure 3**

(a) Right: Temporal variations in photosynthetic photon flux density (PPFD,  $\mu \text{ mol of photons m}^{-2} \text{ s}^{-1}$ ) over a 30-day period (01–31 Mar 2013) measured at every 5-minute intervals 07.00 AM to 05.00 PM. Note that the stochastic variations in PPFD occur in response to the subtle or rapid perturbations in cloud cover conditions that are not captured by a clear sky model. (b) Bottom: Sample images obtained by Total Sky Imager (TSI) capturing cloud cover conditions associated with simultaneously measured PPFD, solar zenith angle (SZA) and the time of the day.



**Figure 4**

Scatterplot-based correlation analysis of the 5-minute PPF<sub>D</sub> (i.e., the objective variable) in respect to the 17 cloud-image derived predictor variables used in training the proposed CSLTM model. Least square regression lines with the coefficient of determination ( $r^2$ ) is included for each sub-panel with the definition of each cloud-image derived predictor variable as per in Table 1.



**Figure 5**

Comparison of the 5-minute PPFD (left axis) plotted for the first 7 days within the CLSTM model's training phase in respect to the 17 cloud-image derived predictor variables. Definition of each predictor (right axis) is as per Table 1.

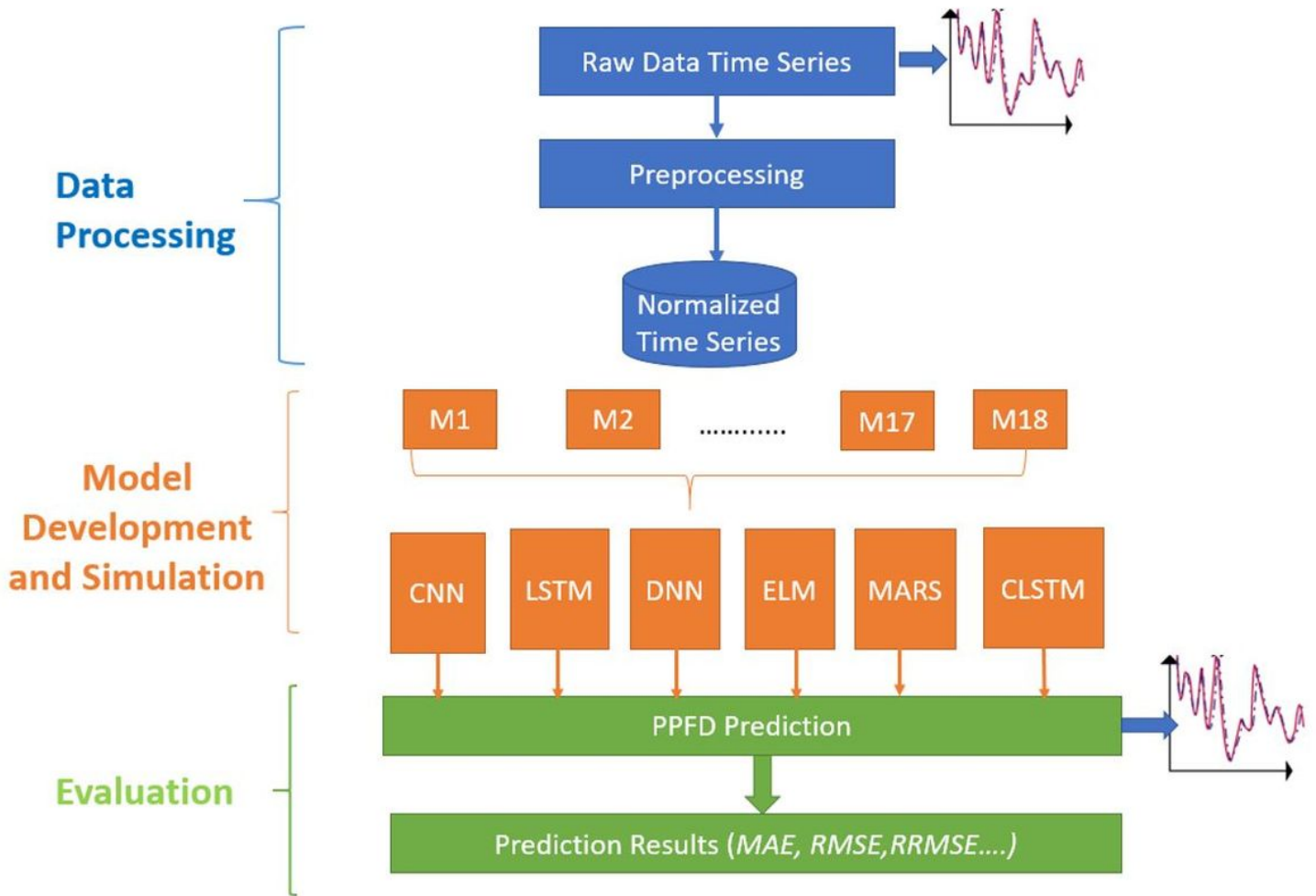
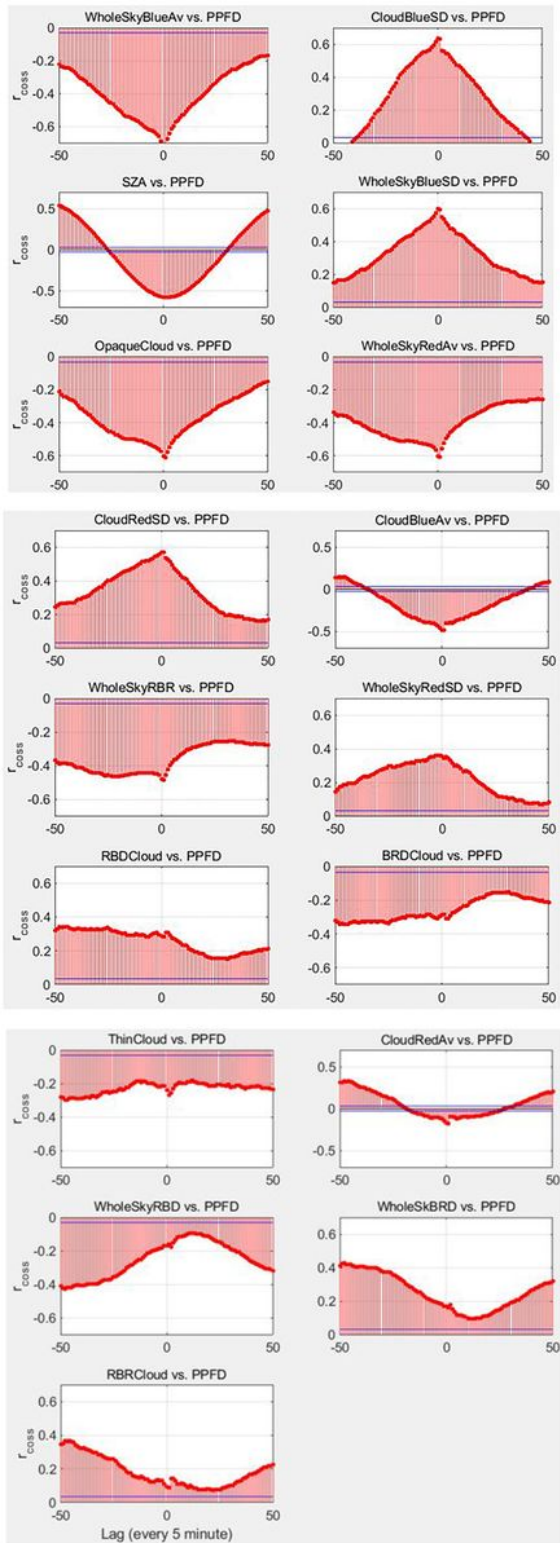


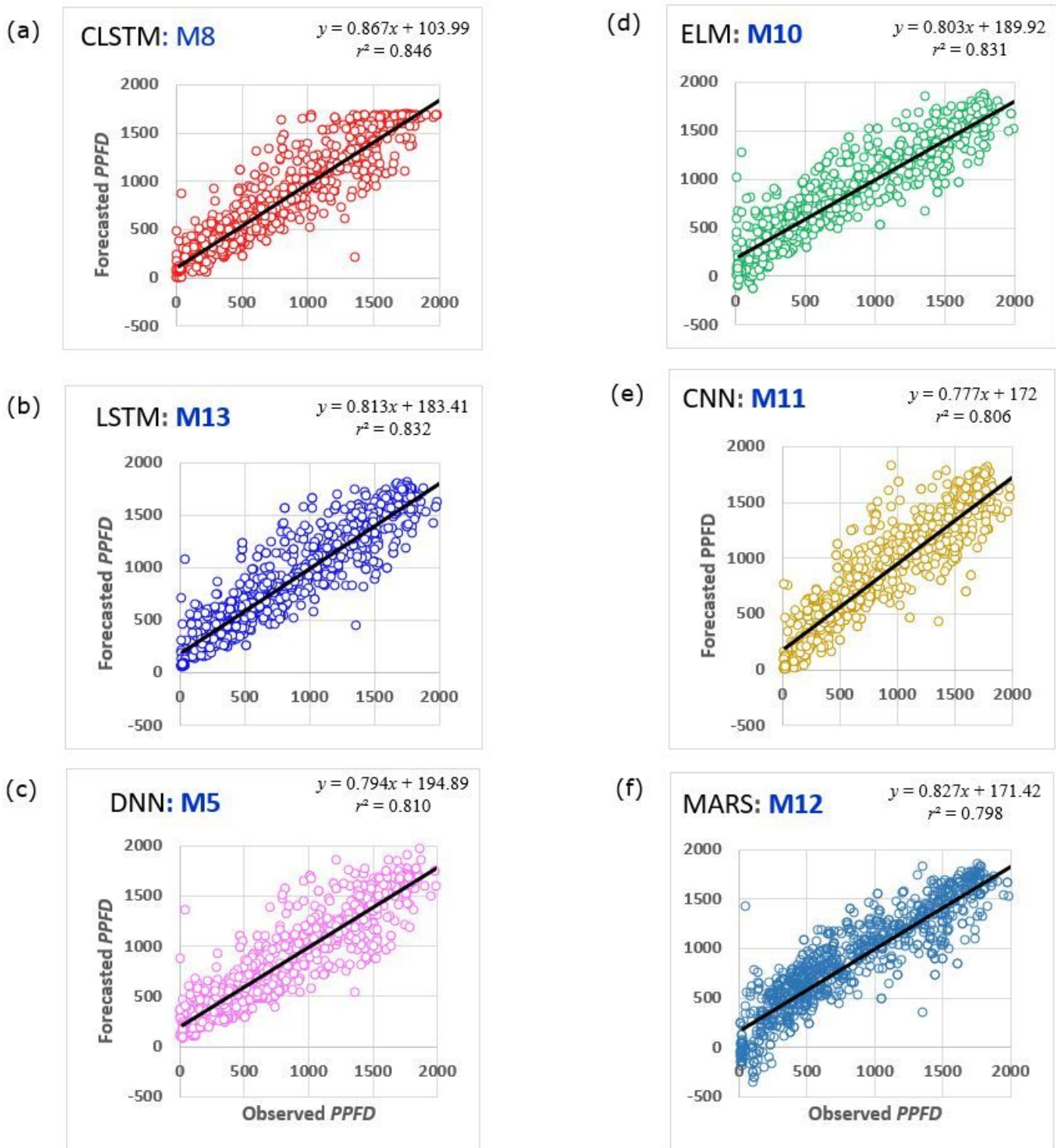
Figure 6

Schematic diagram of the relevant steps in designing the CLSTM predictive model.



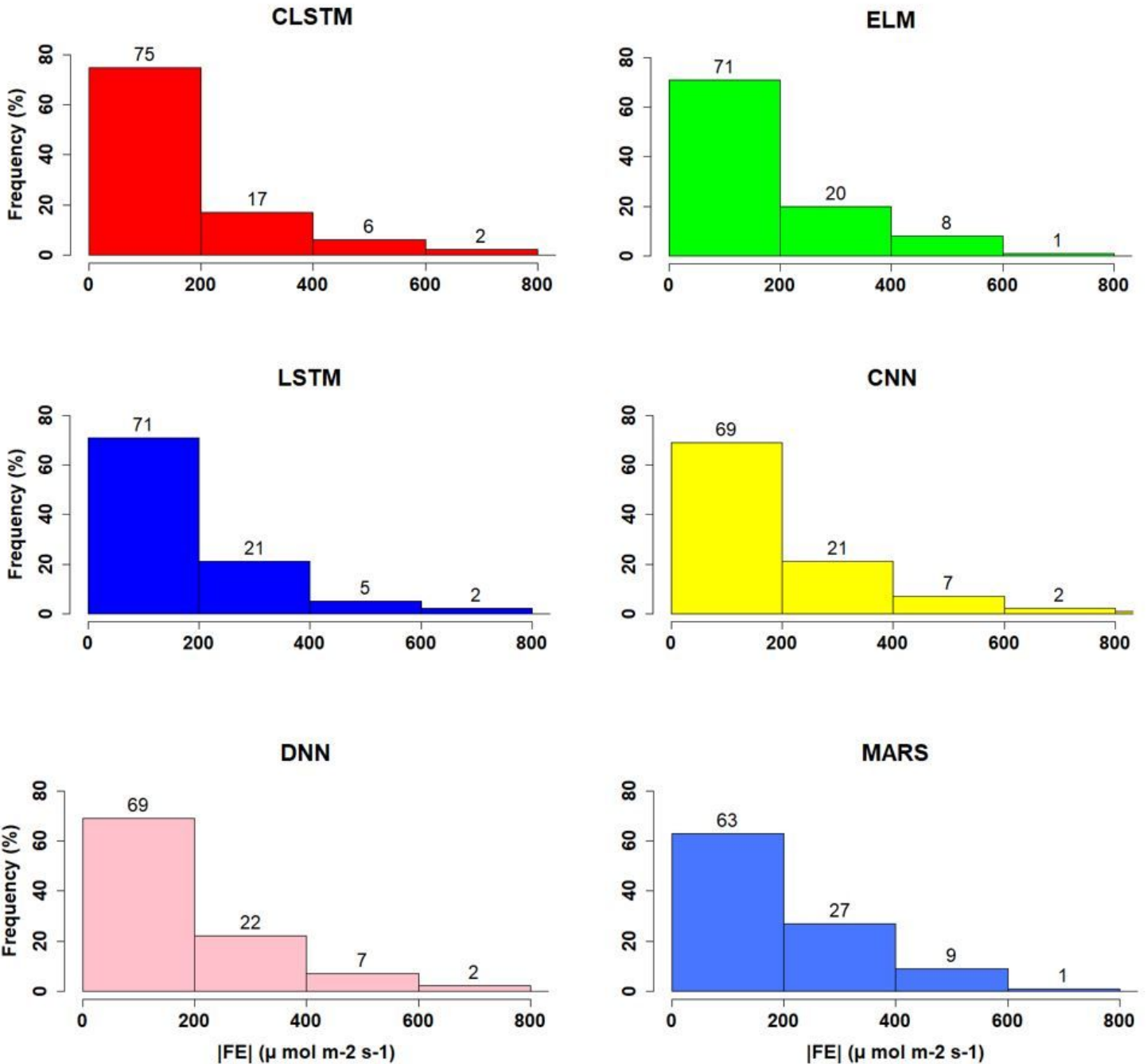
**Figure 7**

Correlograms plotted to identify the degree of covariance between PPFD (i.e., the objective variable) and the 17 different cloud-image derived predictor variables within the CLSTM model's training phase. The y-axis shows cross-correlation coefficient,  $r_{\text{cross}}$  with blue line representing the level at the 95% confidence interval.



**Figure 8**

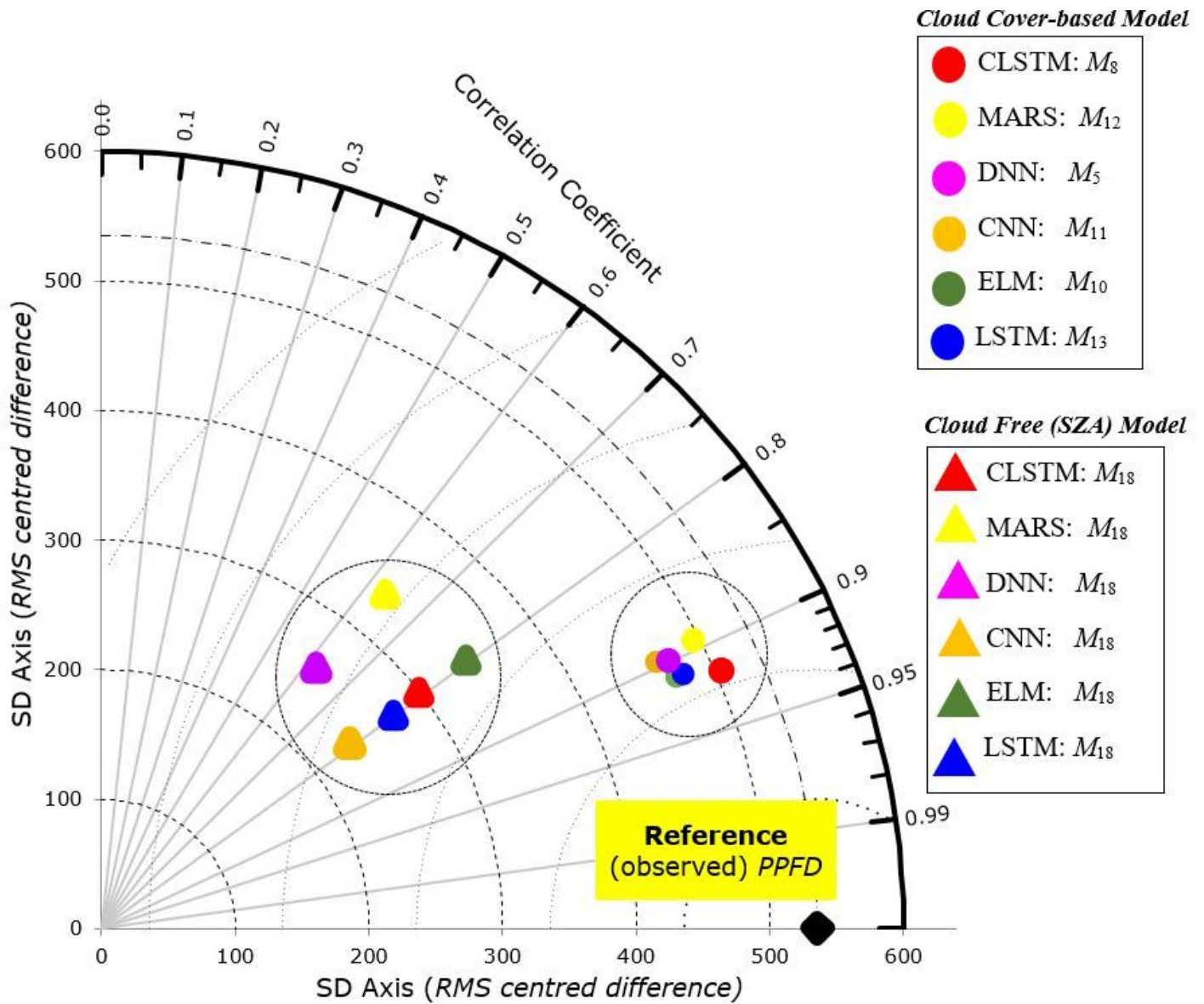
Scatterplots of forecasted against observed PPFD values ( $\mu$  mol of photons  $m^{-2}s^{-1}$ ) emulated by the CLSTM model in the testing phase, compared with benchmark models. Only the optimal results (out of all designated models, M1 to M17) for each predictive algorithm based on best input combinations utilising cloud chromatic statistics and SZA as predictors, as per Table 2, are shown.



**Figure 9**

The percentage frequency of the forecasted error generated by the CLSTM model against the deep learning (i.e., LSTM, CNN, DNN) and machine learning (ELM, MARS)-based models developed using best input combinations utilising cloud chromatic statistics and SZA as the predictors, in accordance with Table 2.





**Figure 10**

Taylor diagram with a concise statistical summary of how well the simulations from the CLSTM predictive model match with the other models in terms of their correlations between observed and forecasted PPFDF, root-mean-square difference and the ratio of the variance in testing phase. Only the most optimal model with cloud cover properties (i.e.,  $M_8$ ,  $M_{13}$ ,  $M_{12}$ ,  $M_5$ ,  $M_{11}$  and  $M_{10}$ ) and without cloud properties (i.e.,  $M_{18}$  trained with SZA as input variable) are shown.

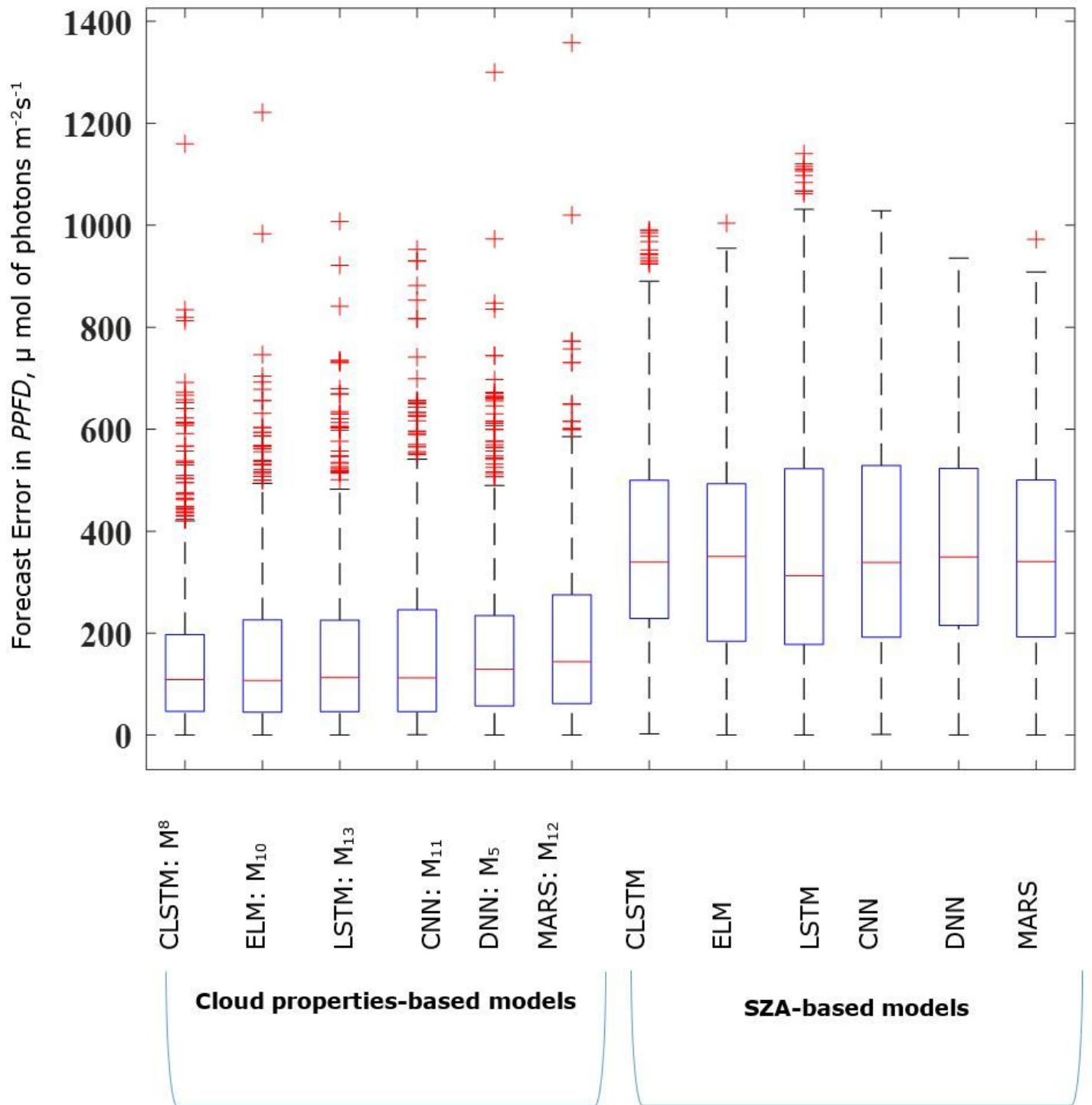


Figure 11

Boxplot of the absolute forecasted error in PPFD:  $|FE| = |PPFD_{for} - PPFD_{obs}|$  within the testing phase using the cloud cover-based and the SZA only reference models. Figure legend should also indicate what the line, box, whiskers and points represent.

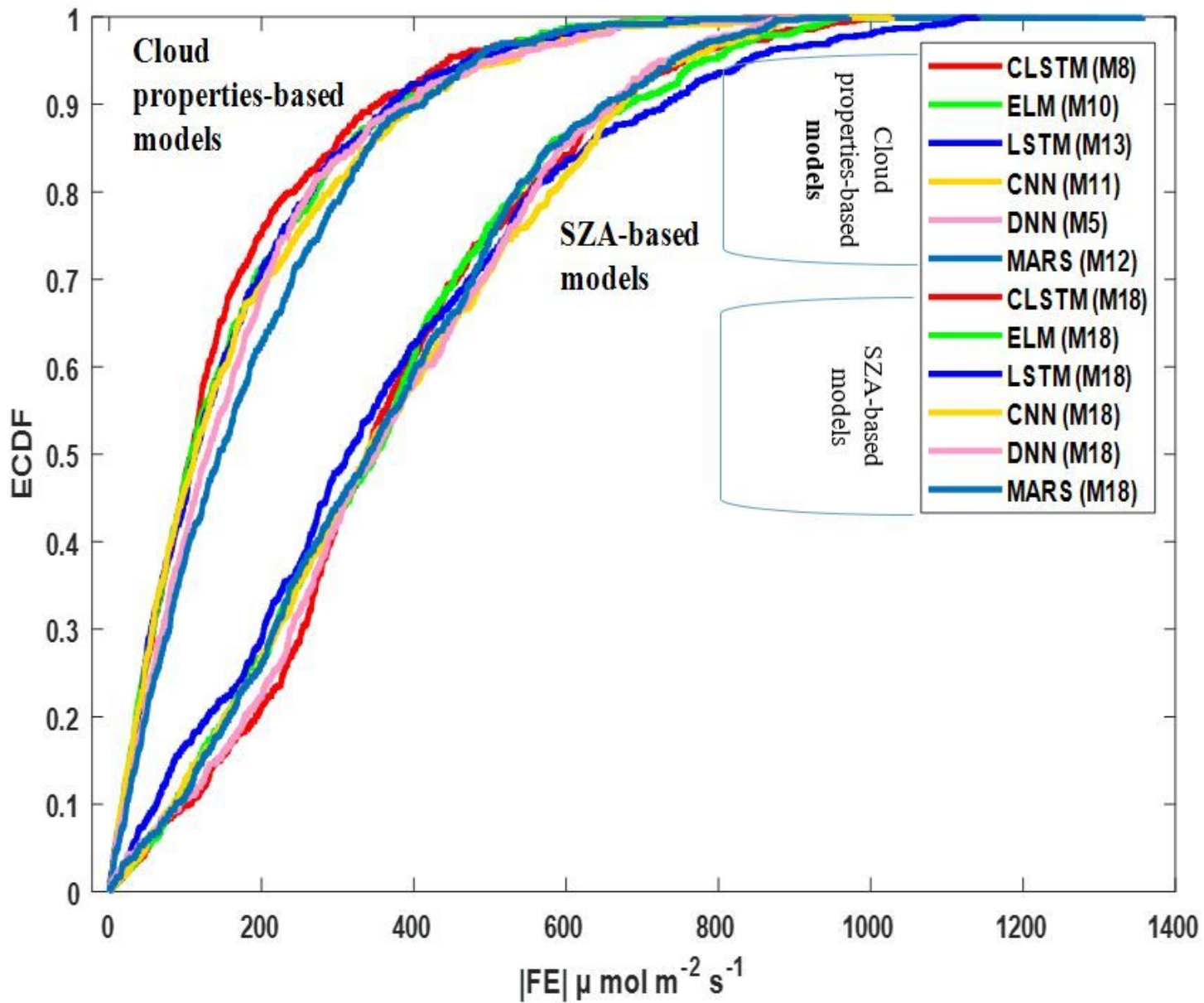
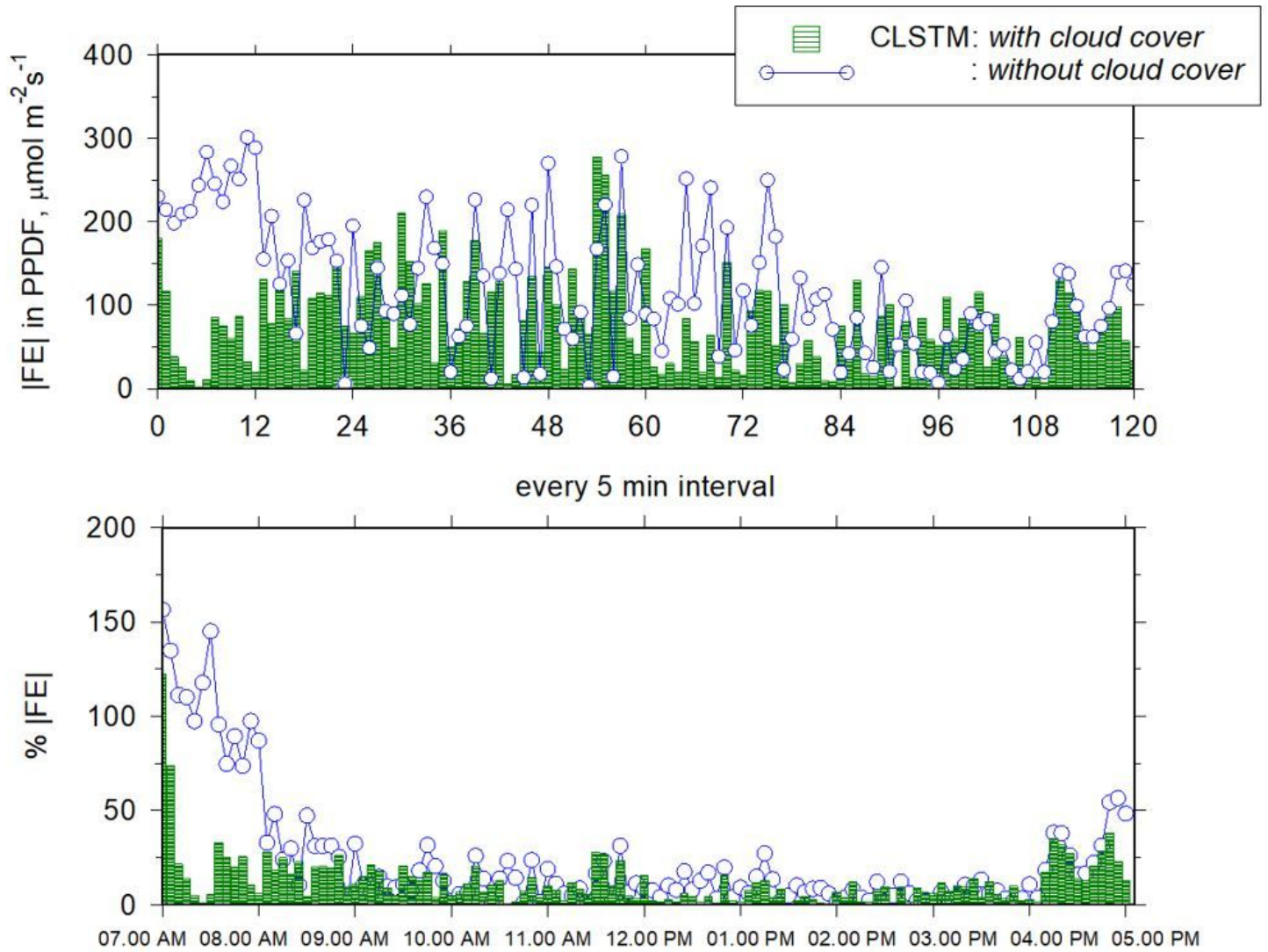


Figure 12

Empirical cumulative distribution function (ECDF) of the PPFD forecasting error  $|FE|$  in the testing phase.



**Figure 13**

The effect of cloud cover properties used as inputs for the CLSTM model with 5-minute forecasted PPFD averaged over the entire testing dataset from 07.00 AM to 05.00 PM.